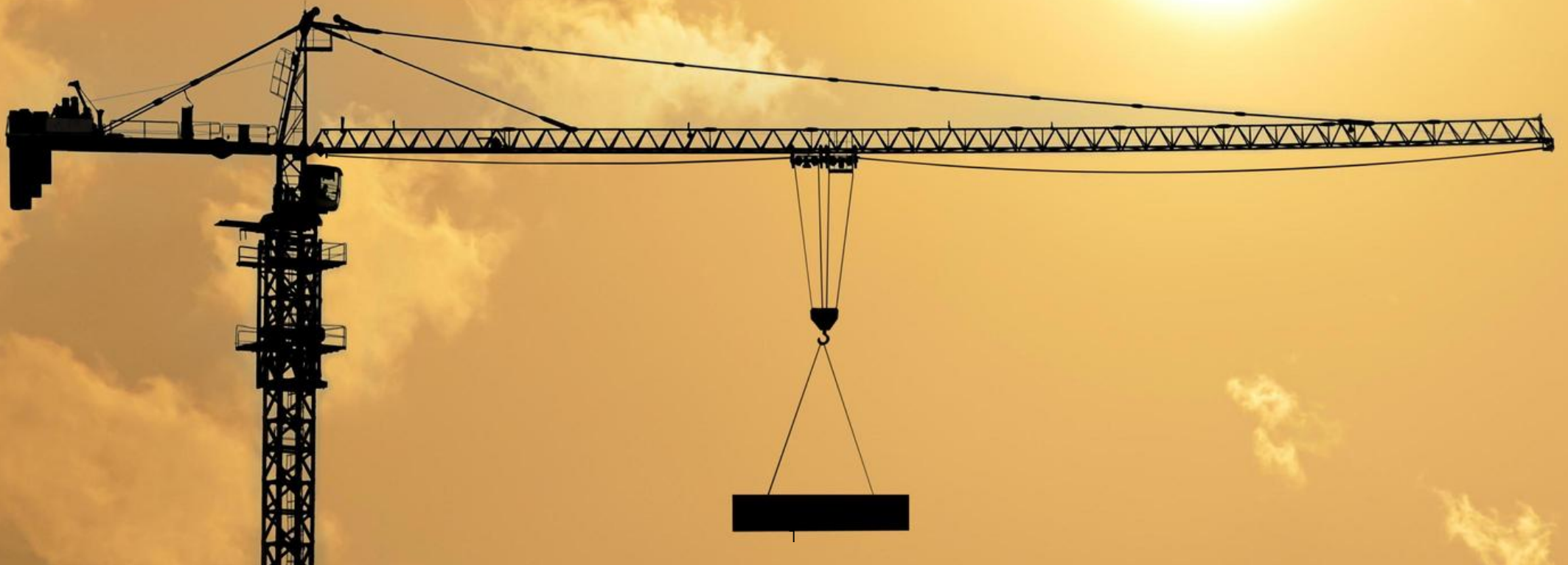


LOFAR 2.0 PIPELINES - LBA COMMISSIONING GROUP



LOFAR 2.0 LBA Commissioning Team

| | |
|--|-------------------------------------|
| Francesco de Gasperin (INAF) | Coordinator, calibration strategies |
| Jort Boxelaar (INAF) | VLBI, calibration strategies |
| Cristina Cordun (ASTRON) | <30 MHz, transients |
| Alexander Drabent (Uni Tautenburg) | Pipeline workflow |
| Henrik Edler (ASTRON) | Calibration strategies |
| Christian Groenveld (Uni Leiden) | <30 MHz |
| Thomas Pasini (INAF) | Extraction |
| Reinout van Weeren (Uni Leiden) | Calibration strategies |
| Marco Iacobelli & Tim Shimwell (ASTRON) | Coordination with other groups |

Other related commissioning groups:

- Pipelines: VLBI
- Pipelines: Pre-process/LINC/Rapthor
- Operations: scheduling

What can we currently do?

- We can do (deep) imaging at 42-66 MHz using Dutch array
- We can do imaging of a few very bright sources with IS
- We can do imaging at <30 MHz with good ionospheric conditions
- It is unclear if we can do IS on moderately fluxy sources
- We probably cannot do wide-field imaging with IS
- We are severely limited by ionospheric conditions

Aims of the team:

- LBA Long Baselines (> 50 km)
- Wide field calibration at < 30 MHz and low-dec
- Reach thermal noise in sub-optimal ionospheric conditions

- Design observing strategies

Deliverables:

- A set of observing strategies and suggested parameters
- Preprocessing strategy
- Processing strategy

First projects and discussion

1. Cal-cal dual-beam observations (1+1+1 hrs) on 3c196 + 3c295 or 3c380 + 3c196: optimize cal strategy for IS, understand direction dependencies for amplitude & full-Jones effects
2. Observations with NenuFAR (6+6 hrs): clarify complications of integrating NenuFAR in our pipelines and understand the benefit for calibration
3. Calibrator observations to get the SEFD down to 10 MHz

Discussion:

- Further commissioning ideas / did we miss anything?
- Link with other groups (avoid duplication of effort)
- Suggestions on strategies