

Comparative analysis of solar activity



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LOFAR Family Meeting 2025
Institut de Physique du Globe de Paris
Sep 26, 2025



LOFAR-BG In preparation

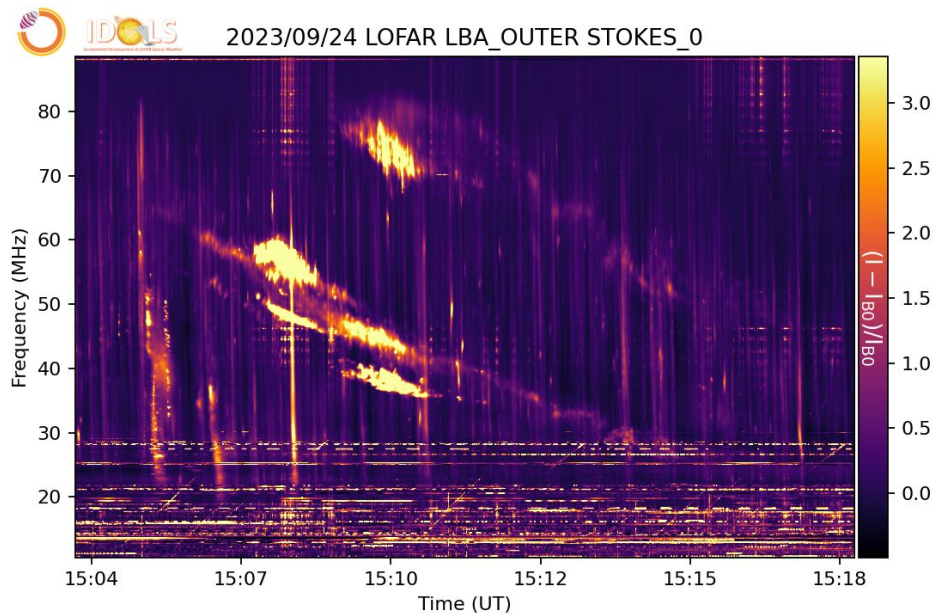
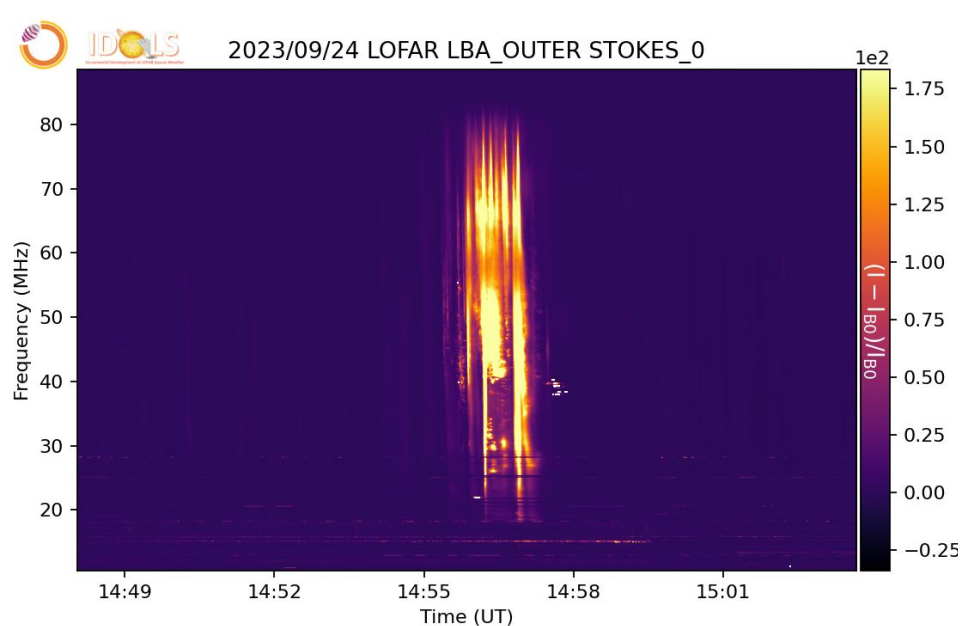


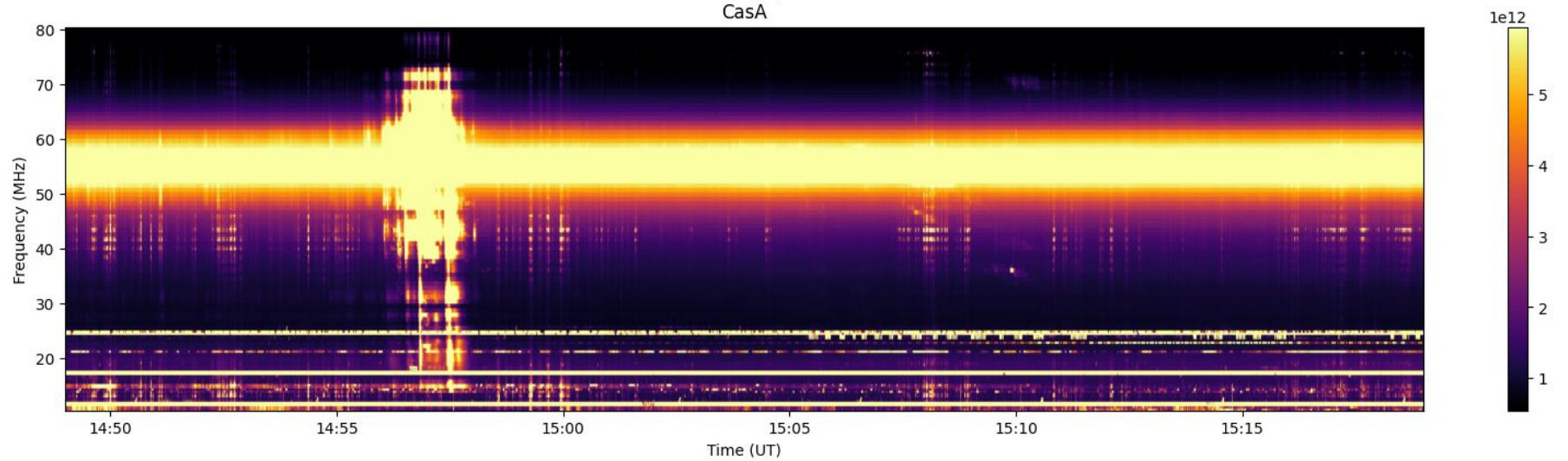
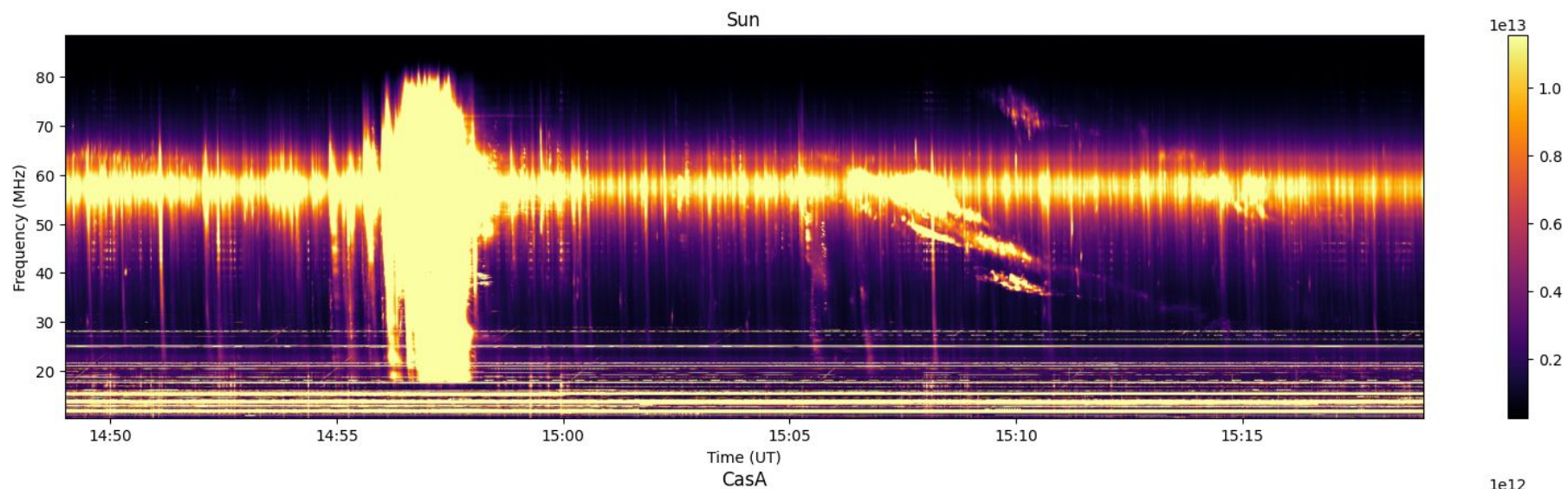
LOFAR Data School 2024



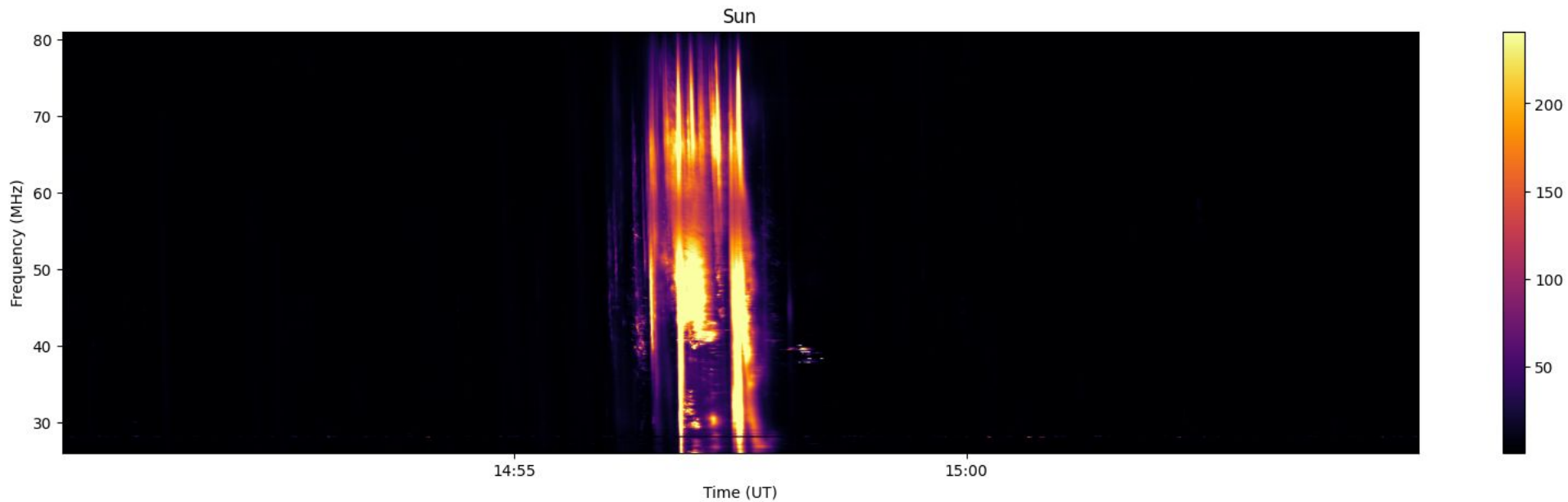
LOFAR LBA IDOLS Beamformed Observations

2023/09/24 14:48-15:18

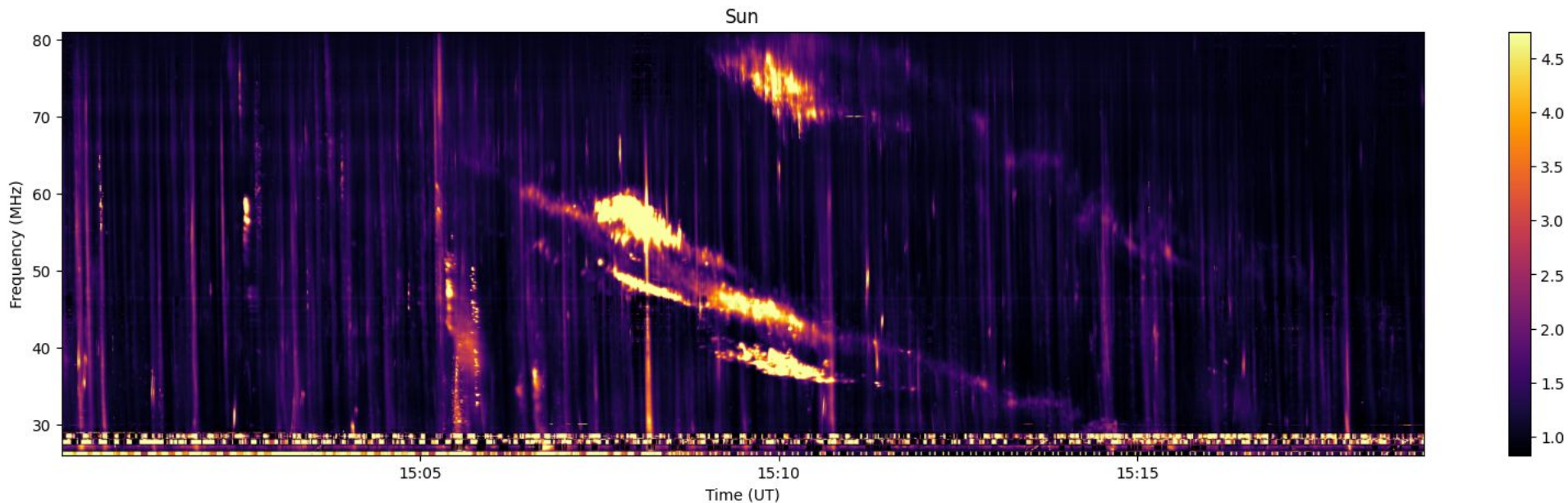




LOFAR LBA 2023/09/24 14:50-15:05 – Type III burst



LOFAR LBA 2023/09/24 15:00-15:20 – Type II burst



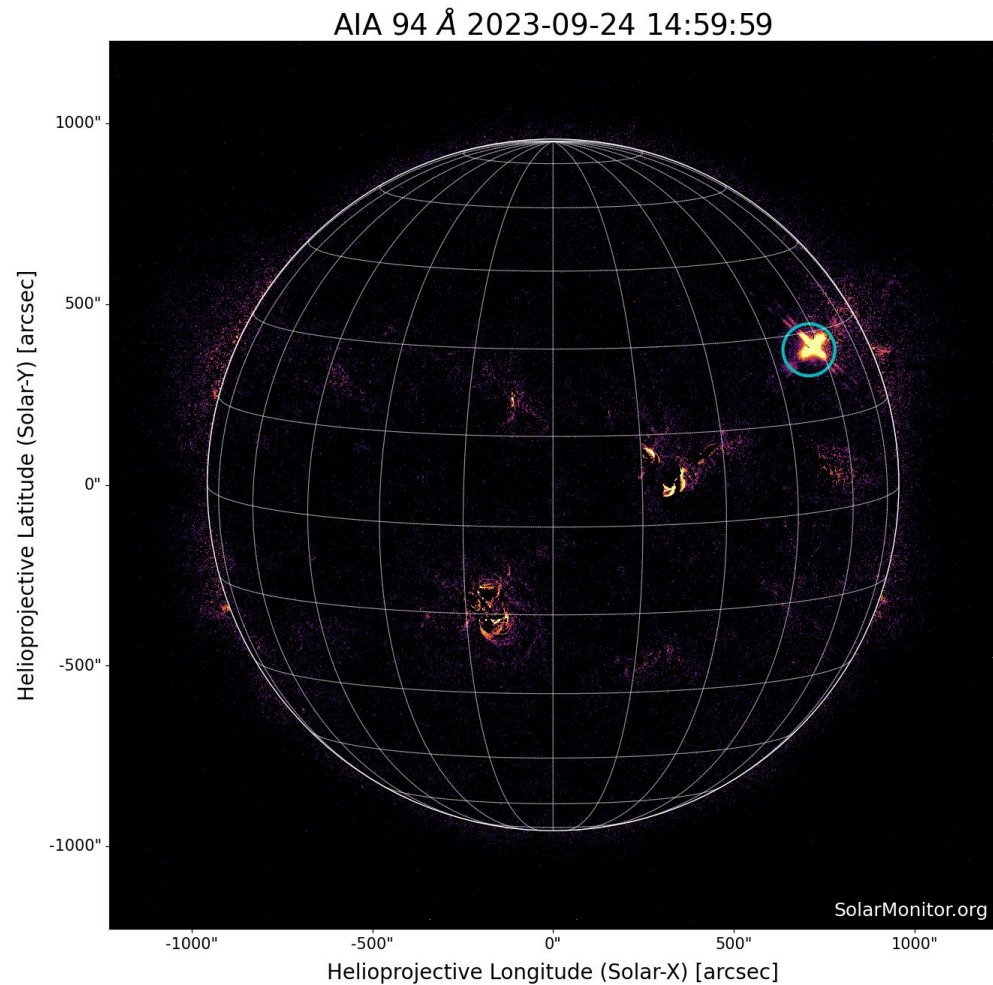
SDO/AIA 2023/09/24

GOES Flare Class: *M1.0*

Active Region:

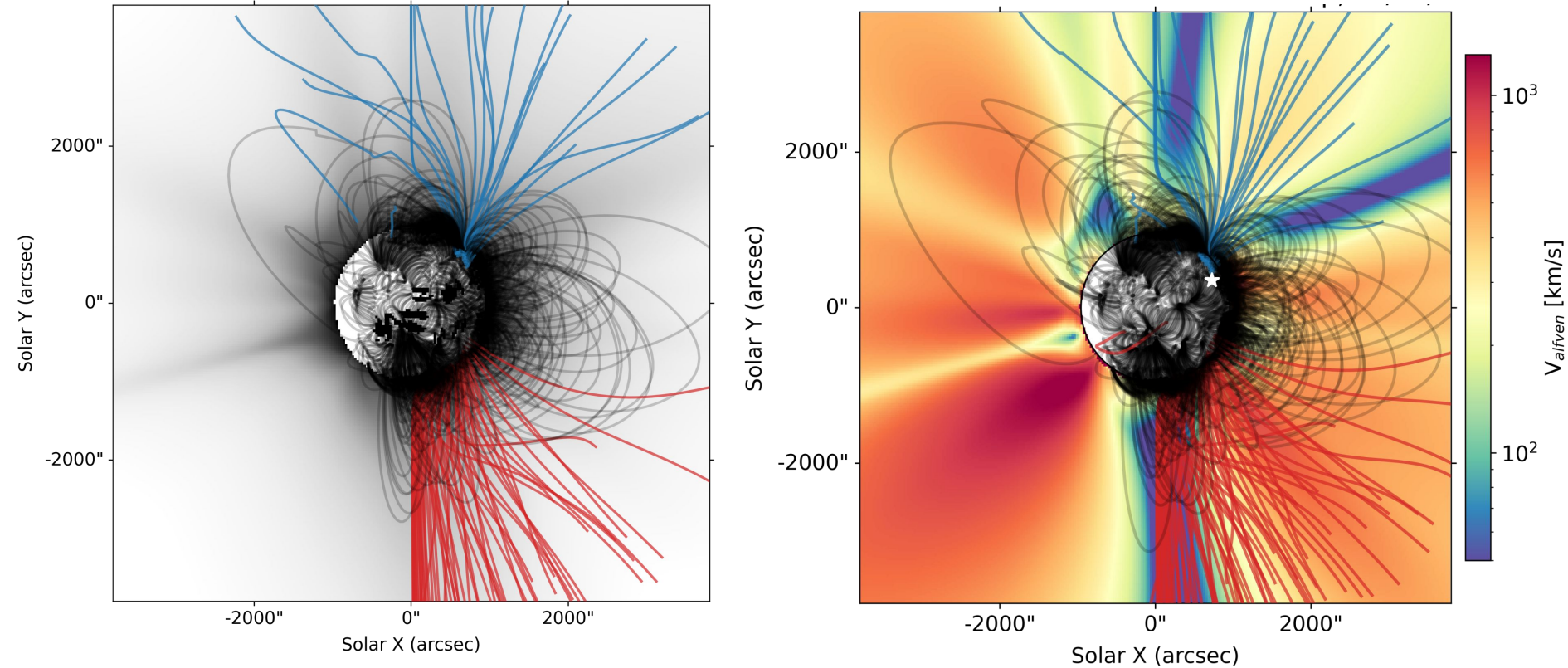
Flare Location (degrees): 56, 27

Flare Location: 729.6", 345.6"



<https://solmon.dias.ie/flare?date=2023092414&id=FL-SSWL-20230924145100>

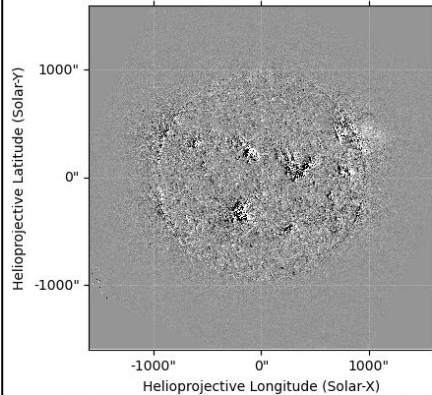
PFSS fields overlaid on MHD valfen speed map 2023/09/24



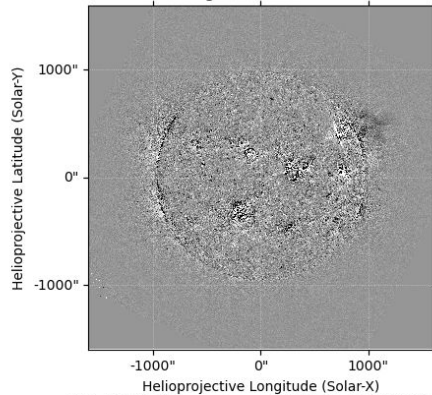
In preparation - Kozarev et al

GOES/SUVI observations 2023/09/24

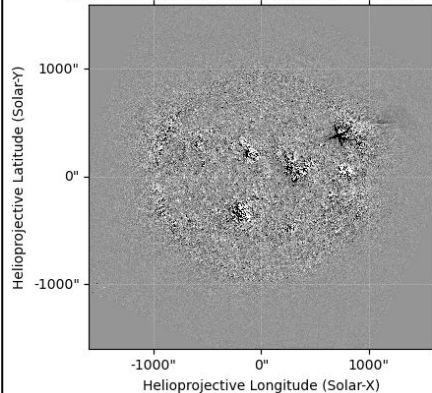
RD, SUVI 284, 2023-09-24T15:00:23



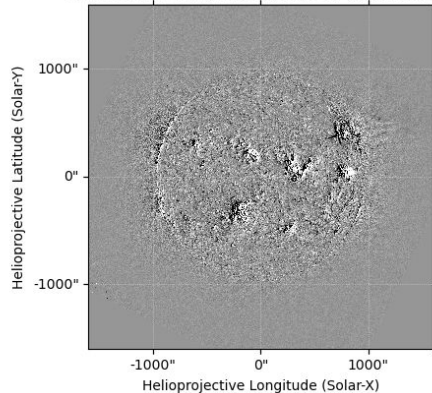
RD, SUVI 284 angstrom, 2023-09-24T15:04:23



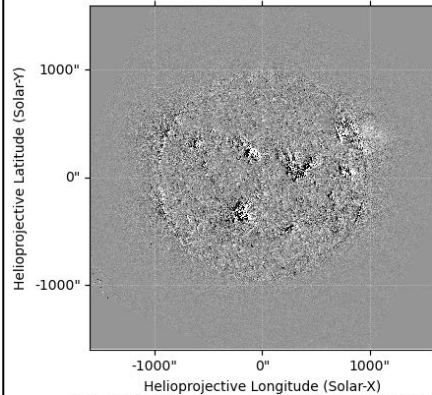
RD, SUVI 284 angstrom, 2023-09-24T15:08:23



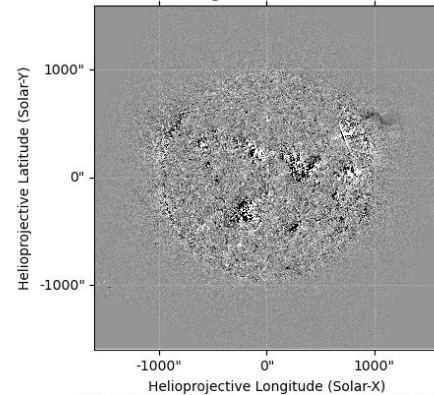
RD, SUVI 284 angstrom, 2023-09-24T15:12:24



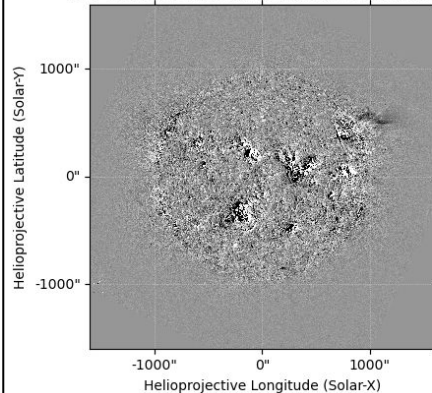
BD, SUVI 284, 2023-09-24T15:00:23



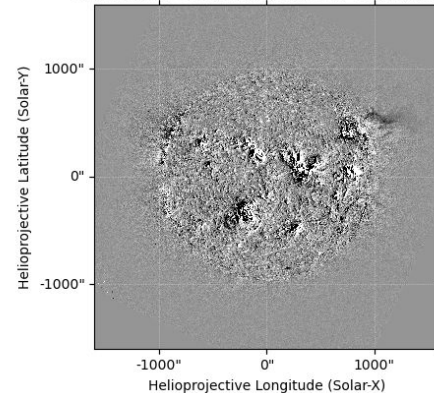
BD, SUVI 284 angstrom, 2023-09-24T15:04:23



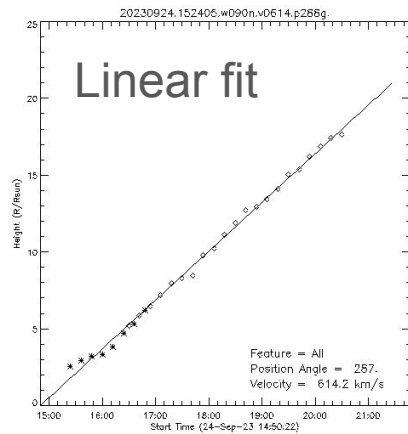
BD, SUVI 284 angstrom, 2023-09-24T15:08:23



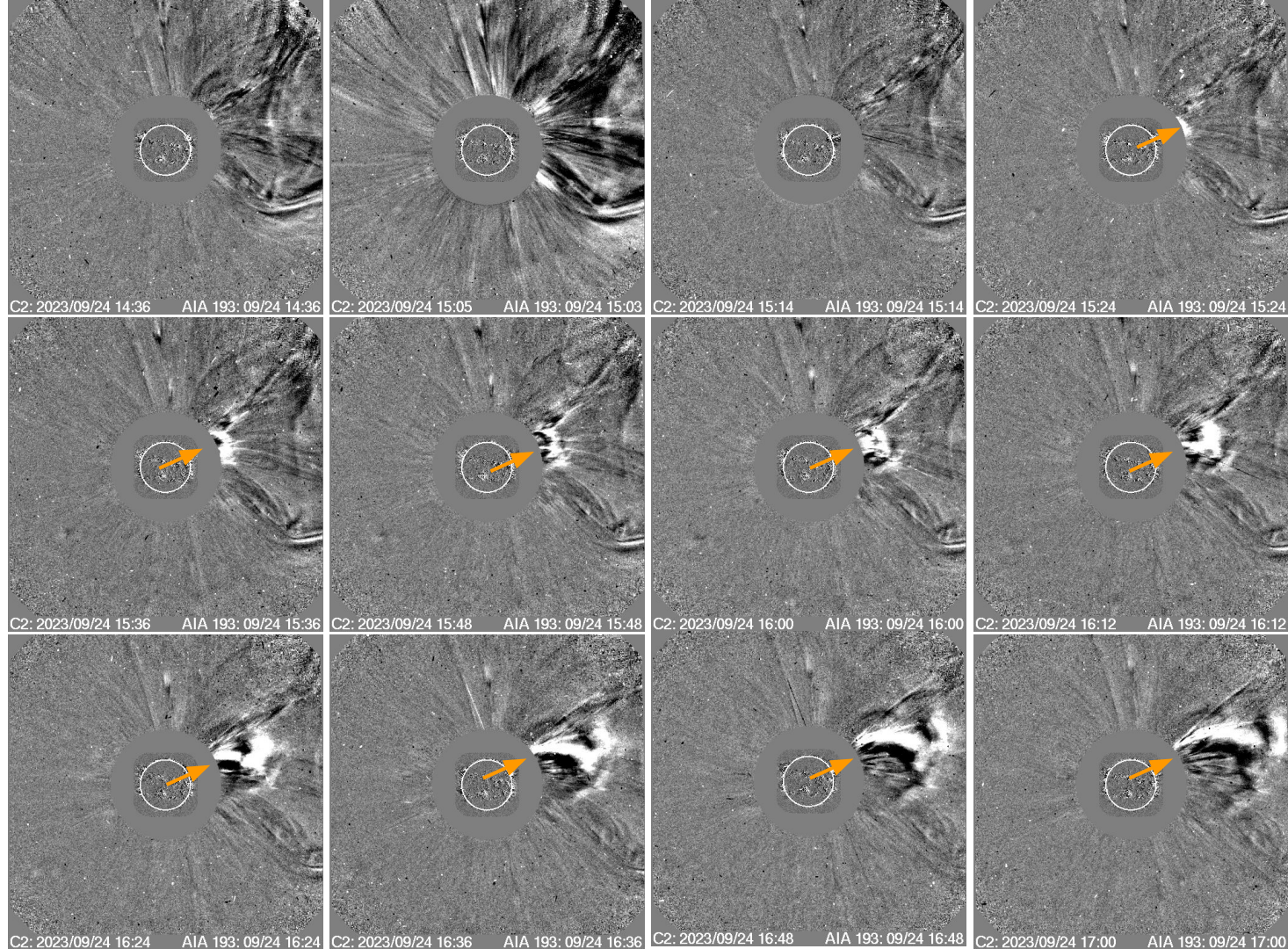
BD, SUVI 284 angstrom, 2023-09-24T15:12:24



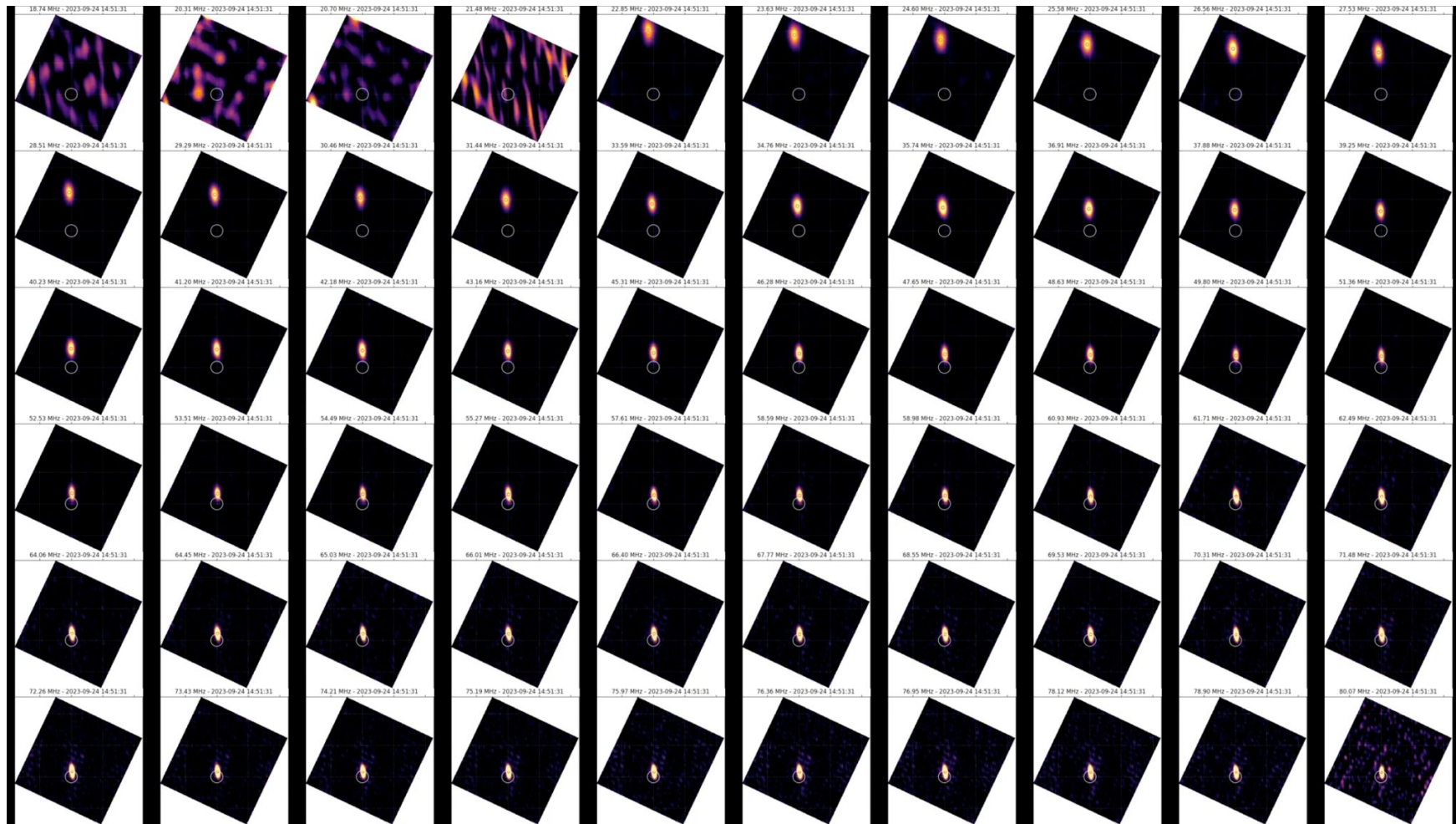
LASCO observations 2023/09/24



LASCO C2 VIDEO

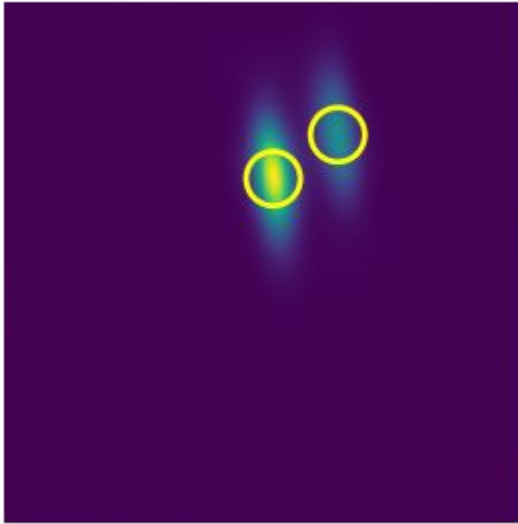


2023/09/24 Interferometry - Significant ionospheric refraction

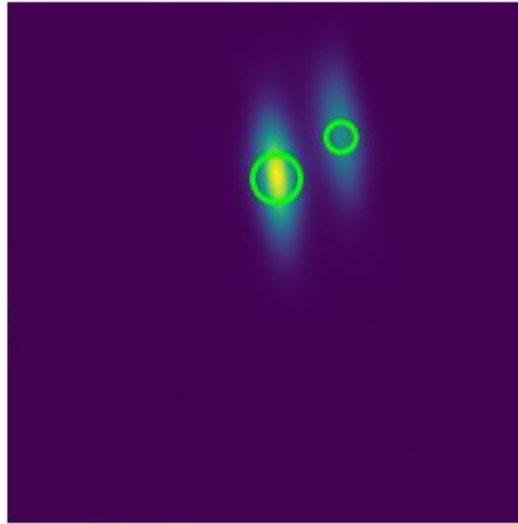


Automatic feature extraction of LOFAR Solar interferometry

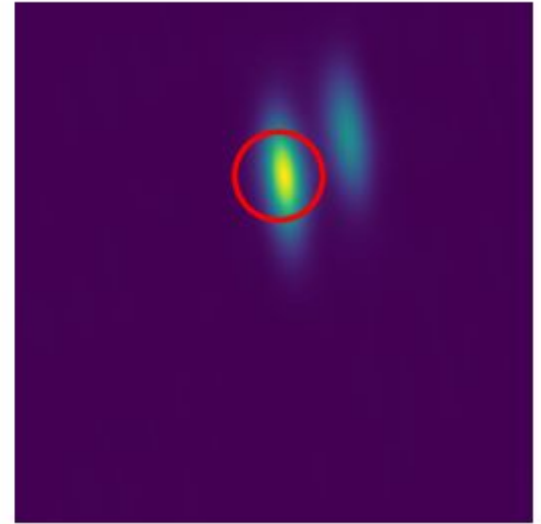
Laplacian of Gaussian



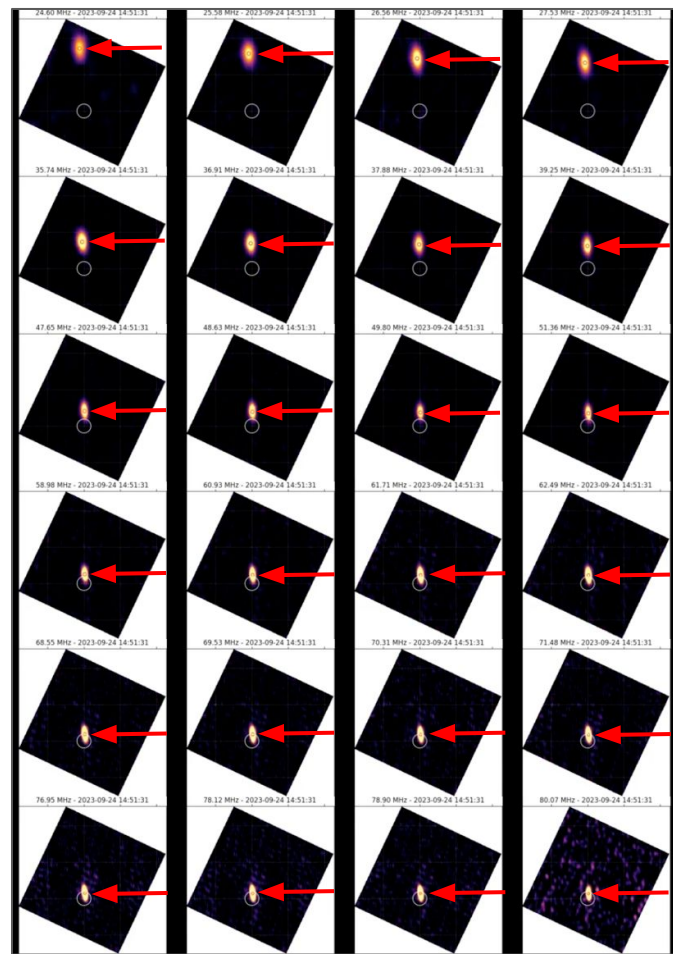
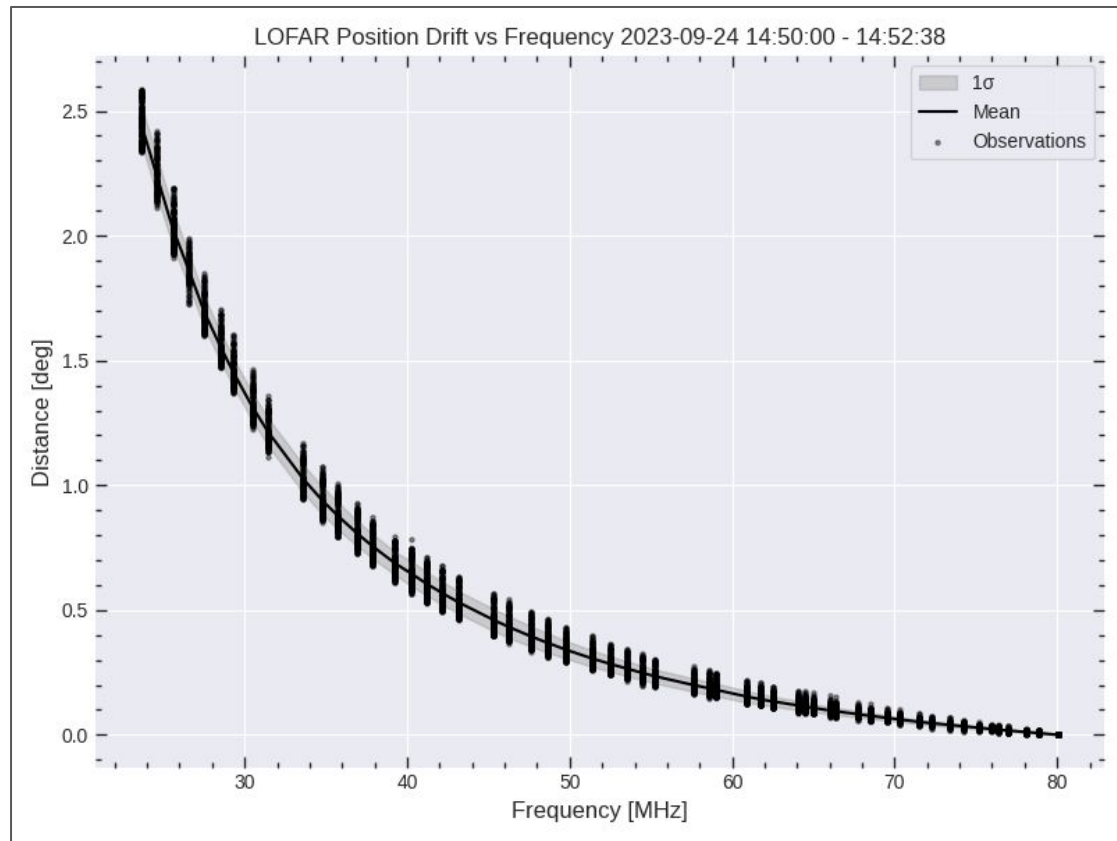
Difference of Gaussian

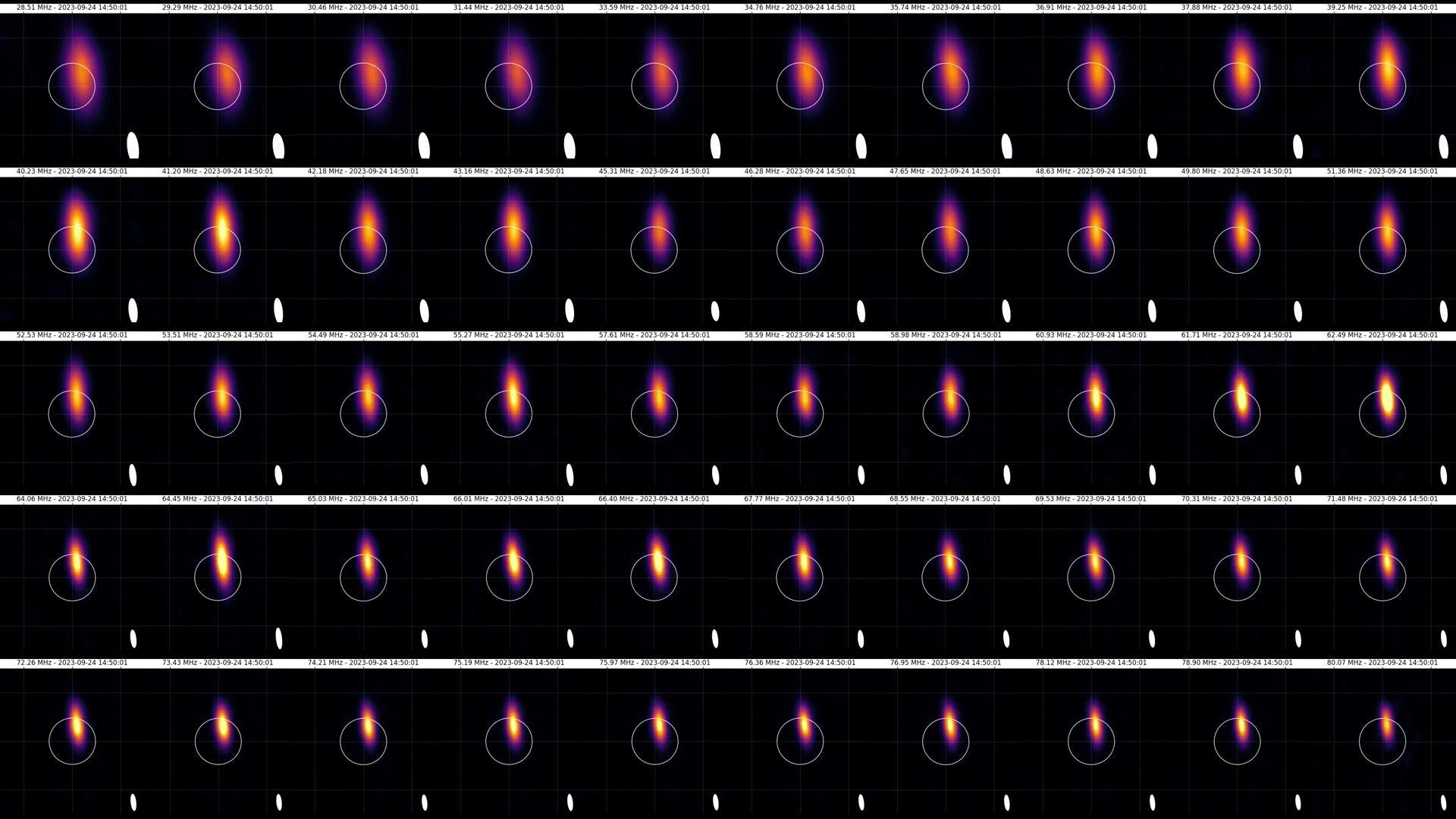


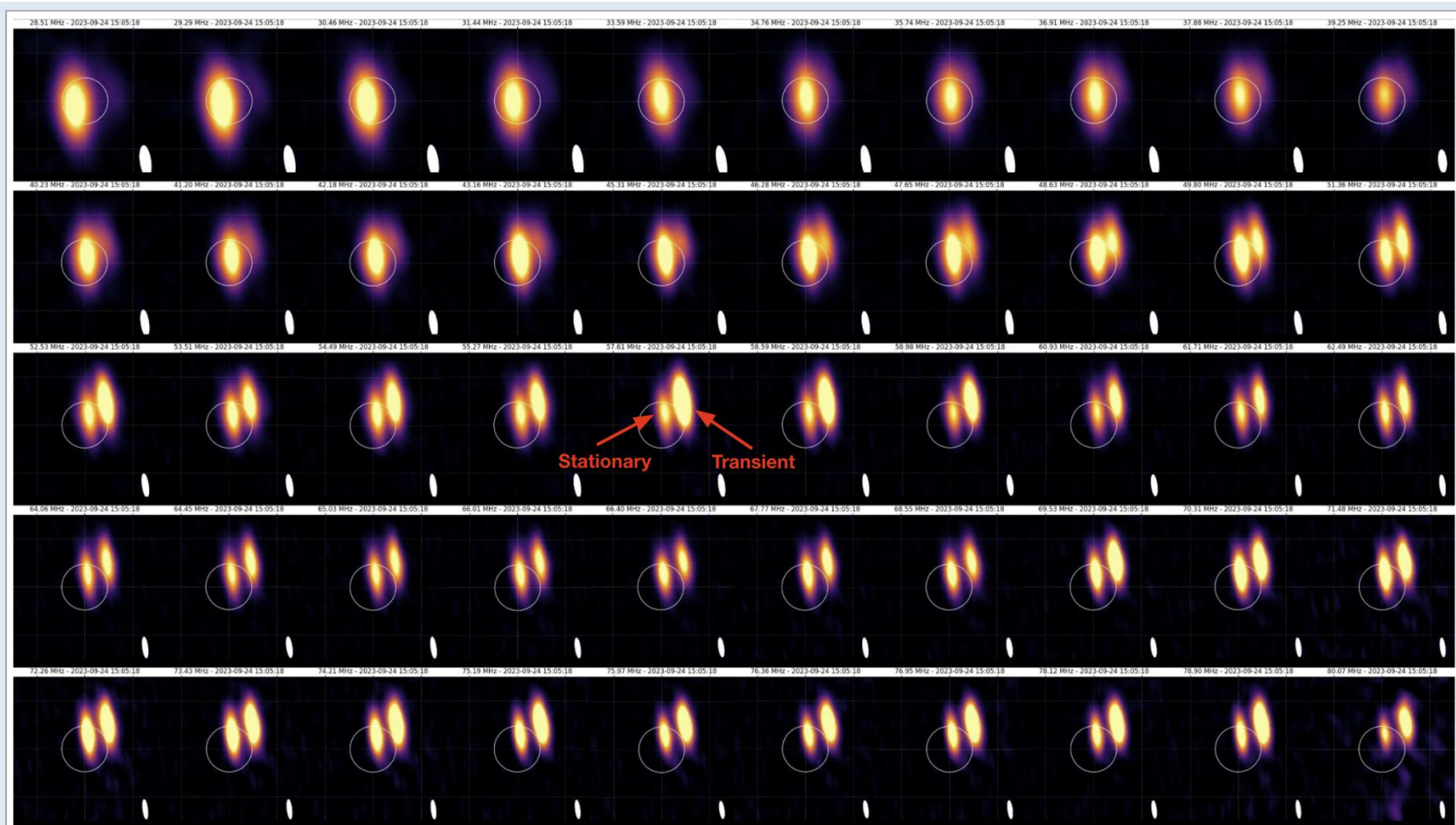
Determinant of Hessian



Position drift of main emission source with frequency

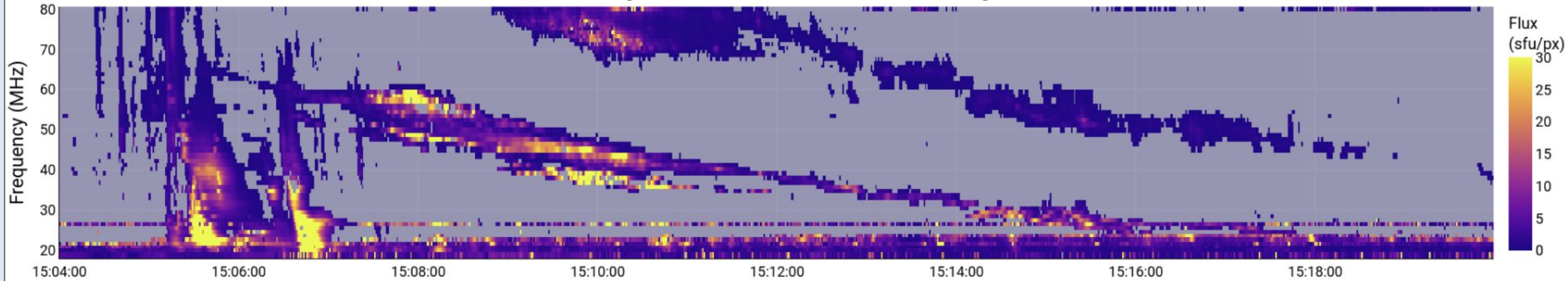




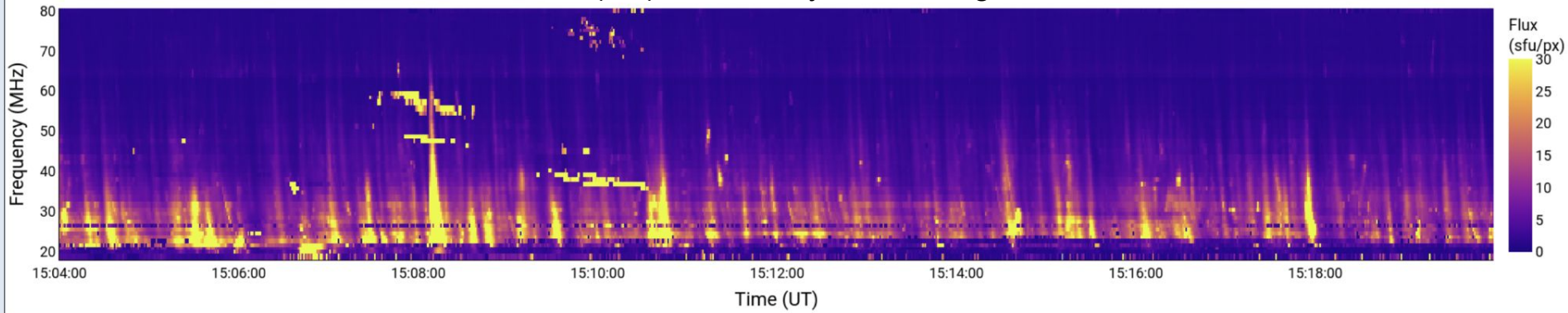


2023/09/24 Extracted sources spectrum

B2 (right) - transient source brightness

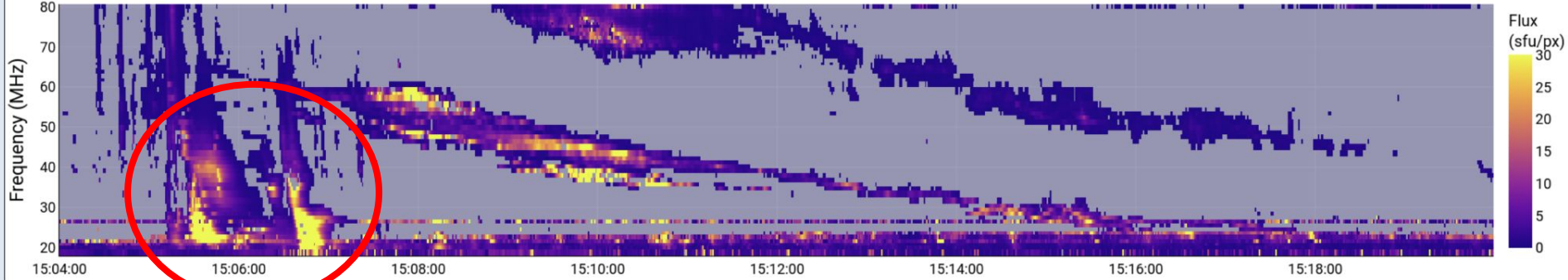


B1 (left) - stationary source brightness

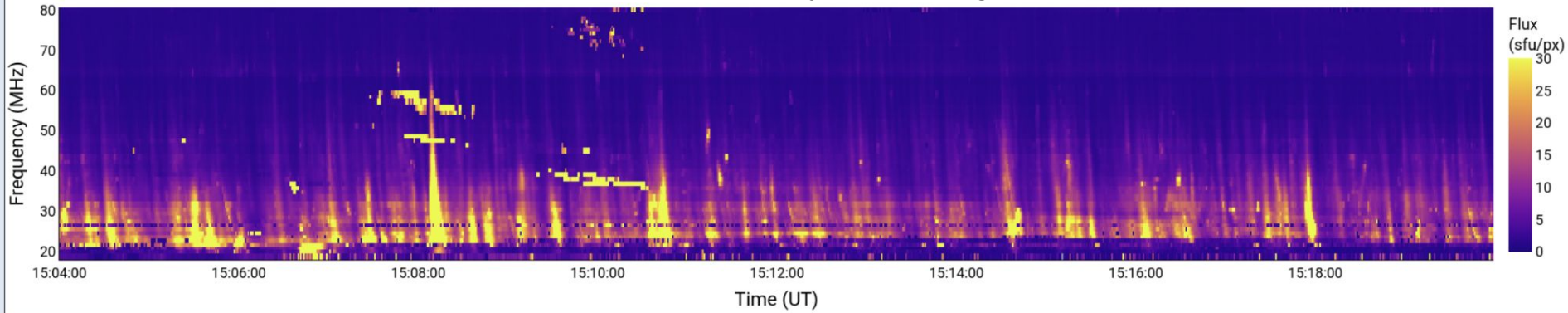


2023/09/24 Extracted sources spectrum

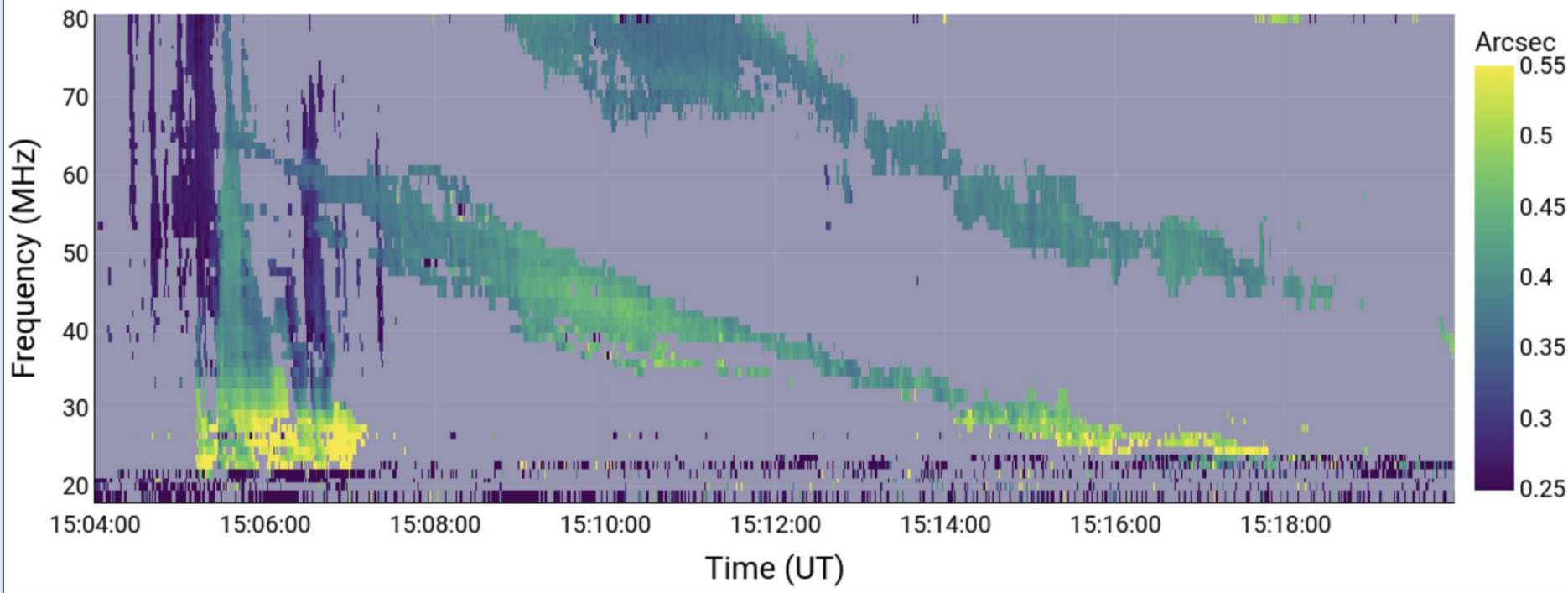
B2 (right) - transient source brightness



B1 (left) - stationary source brightness

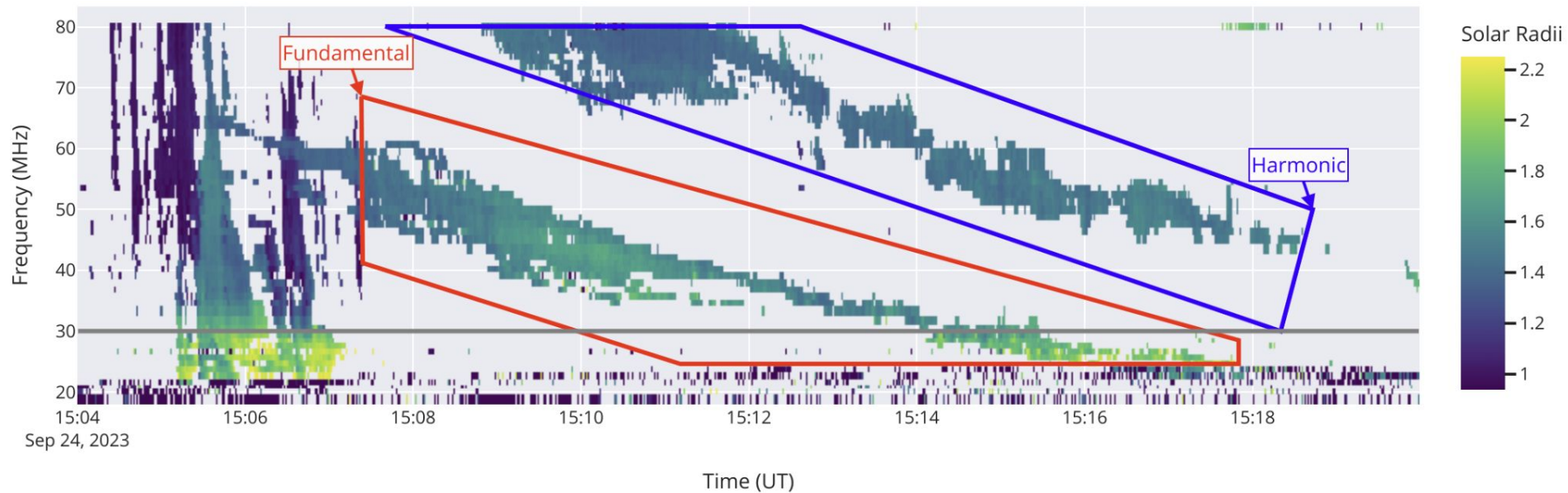


Sources separation



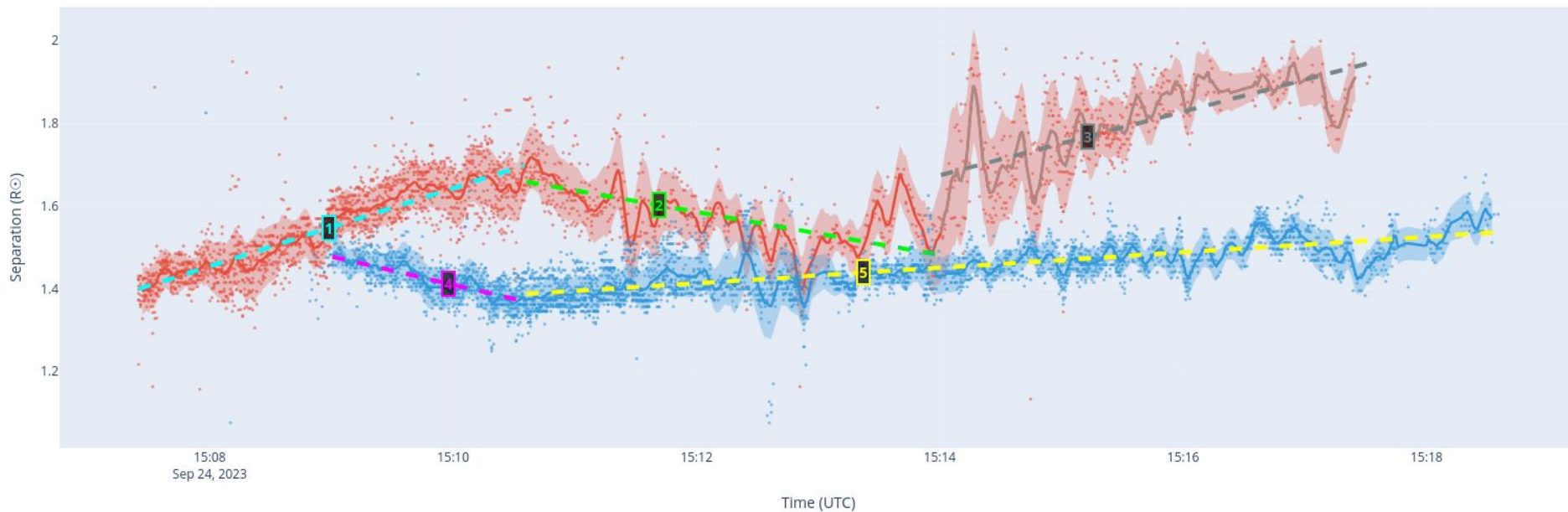
Sources separation analysis

Sources Separation of Type II Emission with Fundamental and Harmonic Regions



Sources separation analysis

Separation Over Time - TypeII Fundamental and Harmonic (Segments 1-5)



Sources

Helioprojective Longitude (Solar-X)
BD, SUVI 284 angstrom, 2023-09-24T15:12:24

Helioprojective Latitude (Solar-Y)

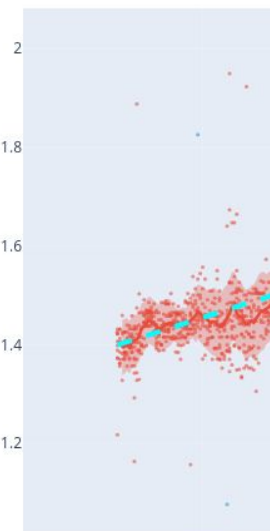
1000"
0"
-1000"

-1000"

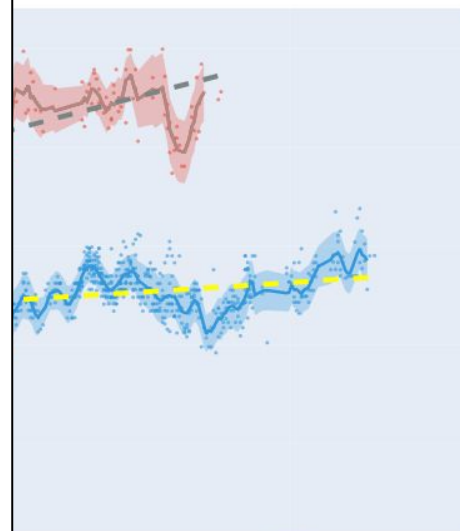
0"

1000"

Helioprojective Longitude (Solar-X)



15:08
Sep 24, 2023



15:18

In preparation -

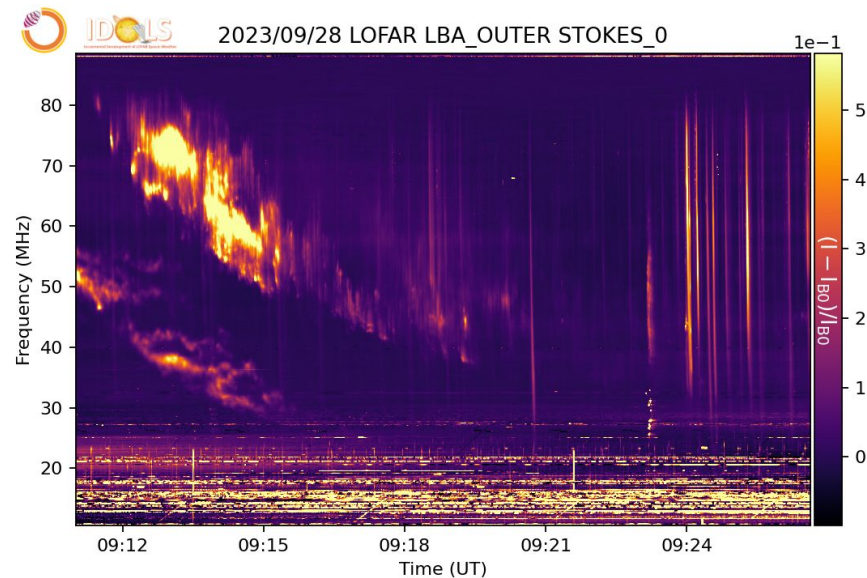
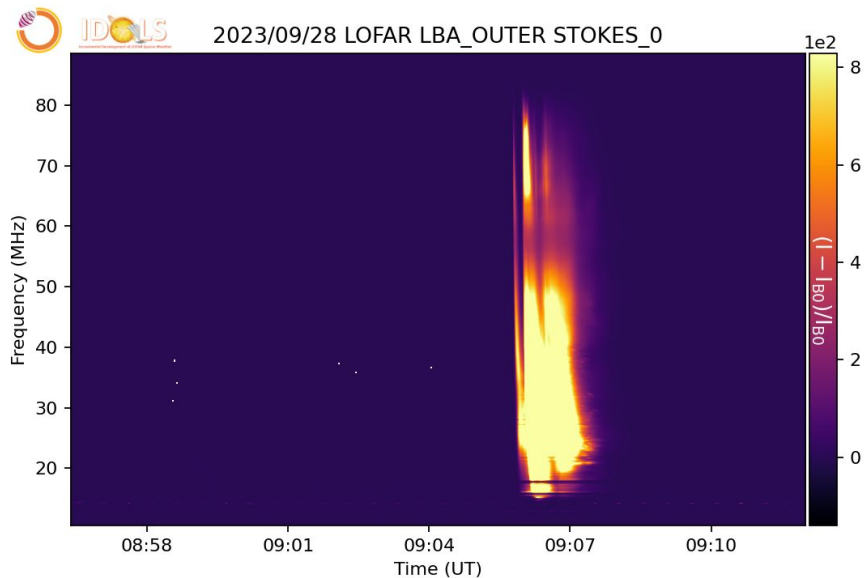
Velocity Segments Analysis

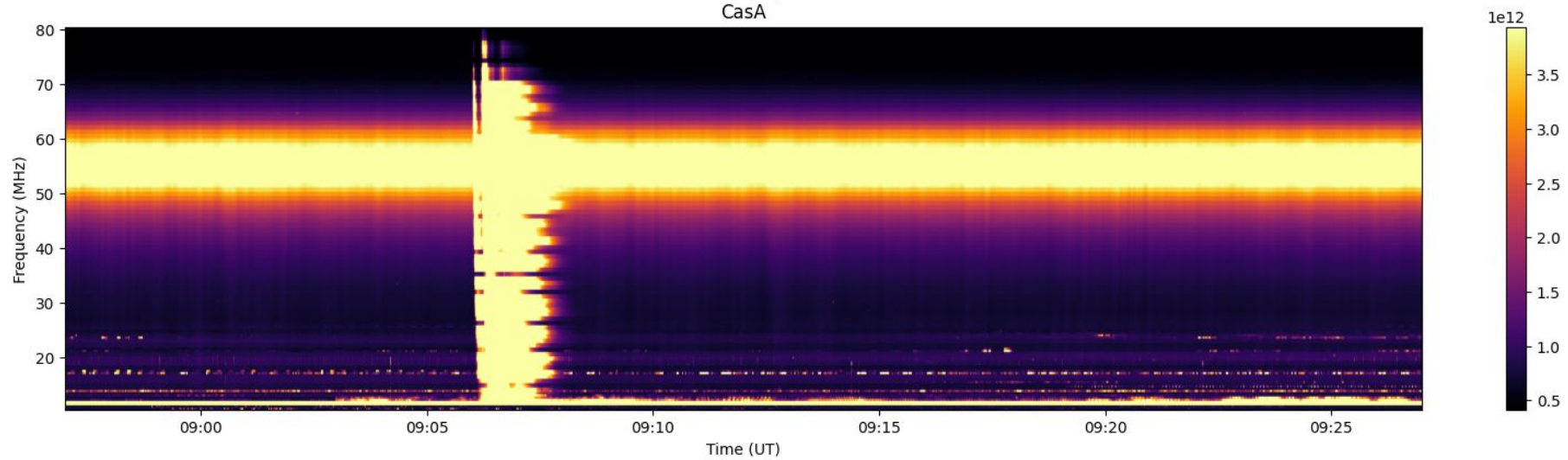
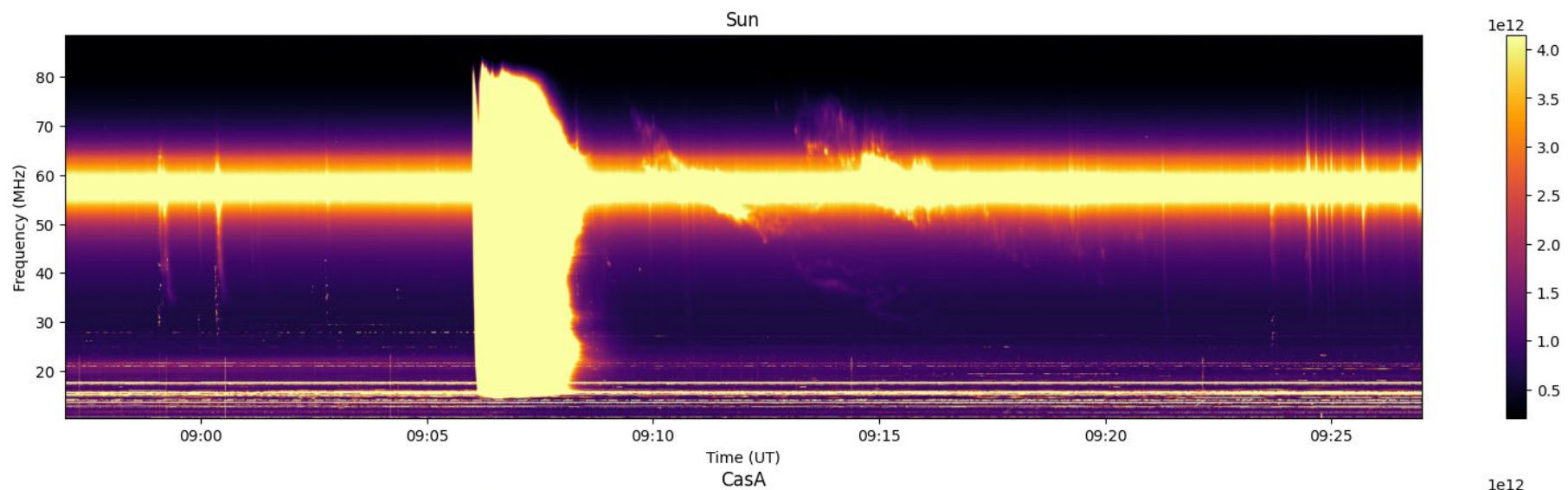
Segment	Time Range (UTC)	Velocity (km/s)	Duration (seconds)	MAD ($R_{\odot}$)	RMS ($R_{\odot}$)	Max Residual ($R_{\odot}$)
1	15:07:24 - 15:10:36	1093.9 ± 311.0	191	0.0425	0.0594	0.4776
2	15:10:36 - 15:14:00	-604.1 ± 404.6	204	0.0592	0.0796	0.3784
3	15:14:00 - 15:17:31	892.2 ± 499.2	212	0.0758	0.1001	0.5990

Segment	Time Range (UTC)	Velocity (km/s)	Duration (seconds)	MAD ($R_{\odot}$)	RMS ($R_{\odot}$)	Max Residual ($R_{\odot}$)
4	15:07:58 - 15:10:36	-796.5 ± 256.1	158	0.0289	0.0445	0.4900
5	15:10:36 - 15:18:35	217.0 ± 98.7	479	0.0340	0.0464	0.3490

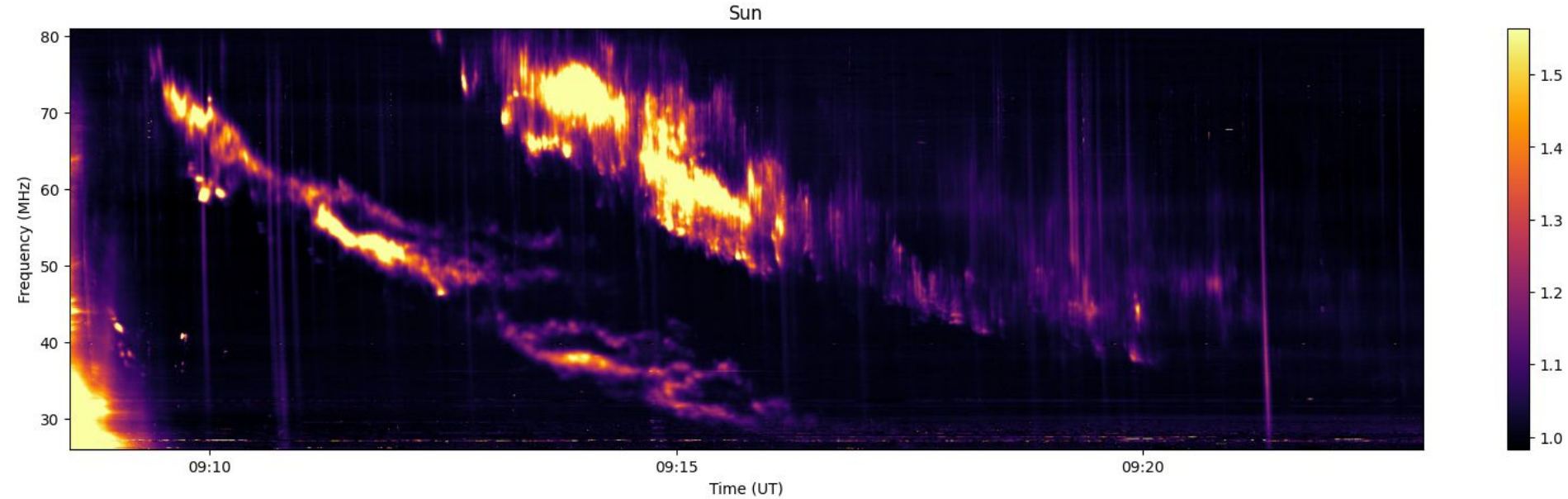
LOFAR LBA IDOLS Beamformed Observations

2023/09/28 08:56-09:26



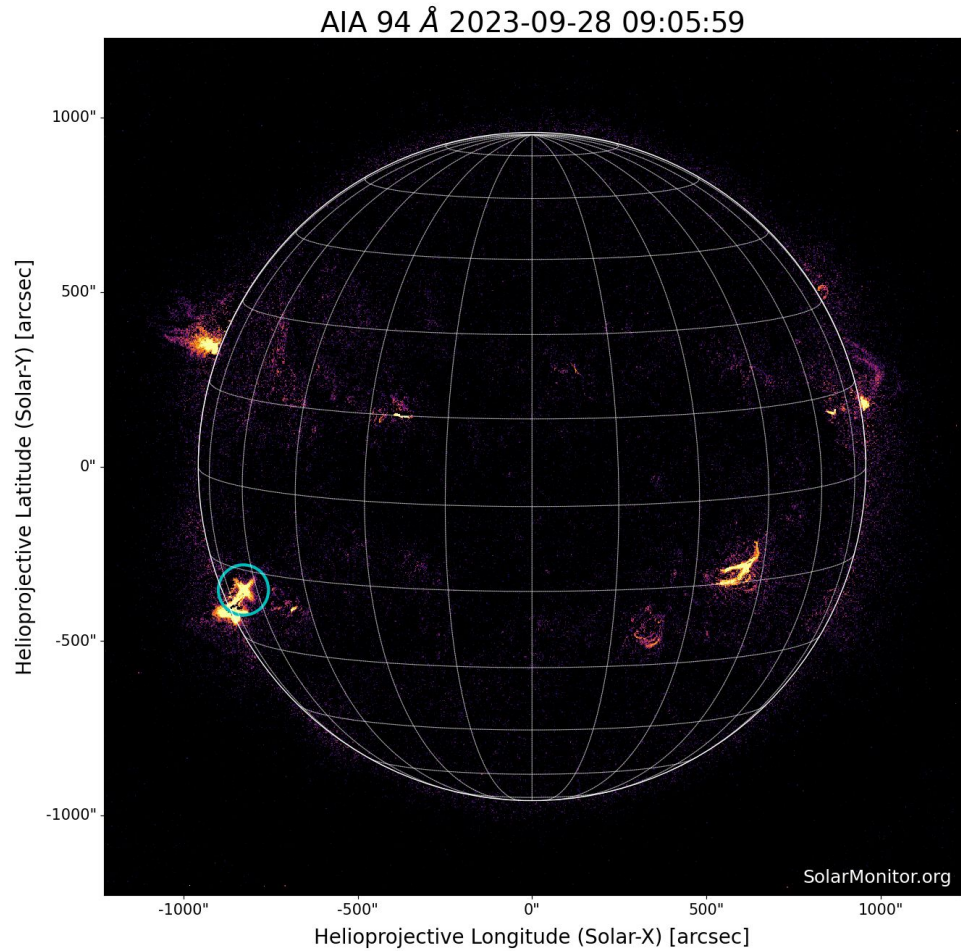


2023/09/28 Type II burst calibrated



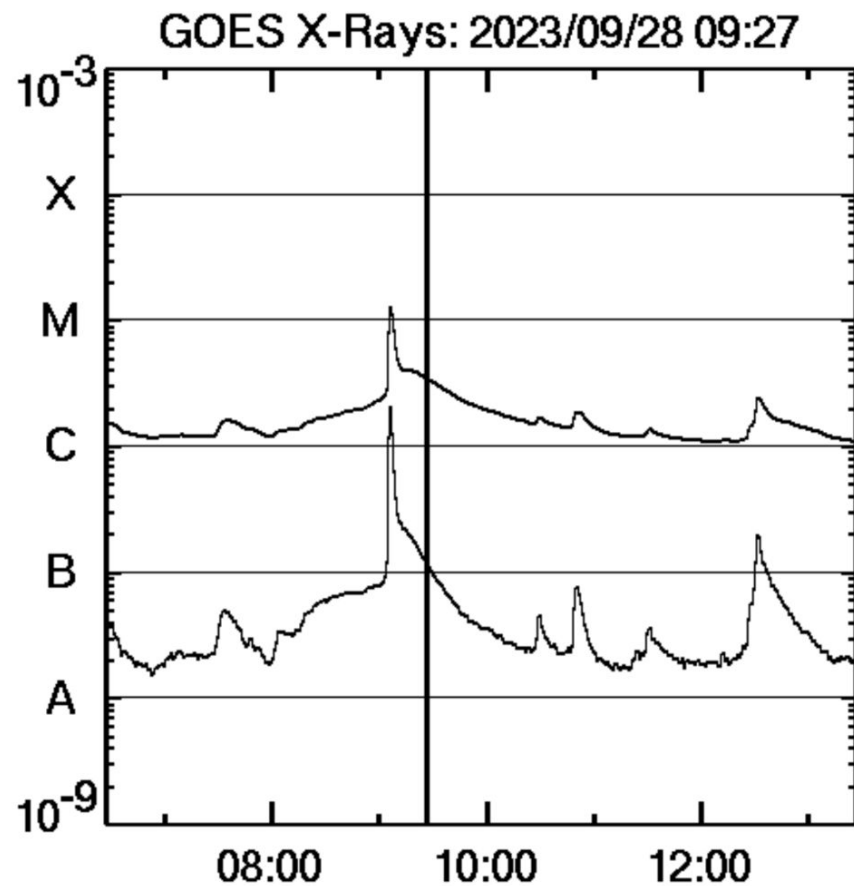
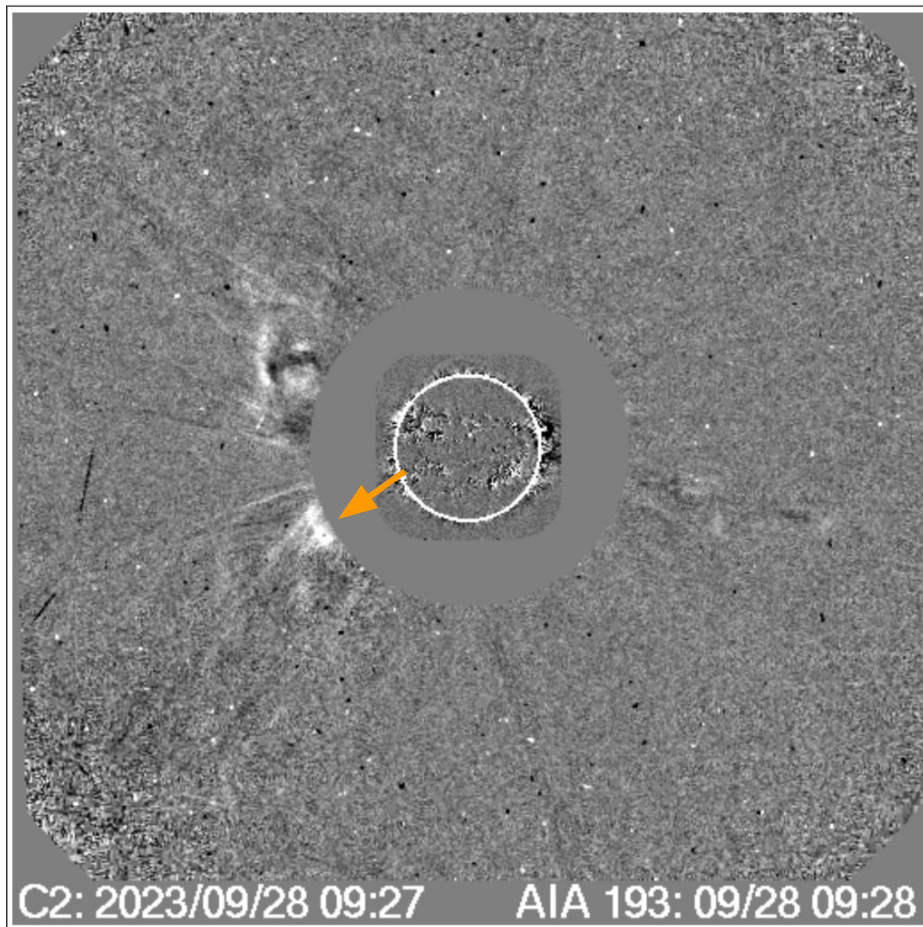
SDO/AIA observations 2023/09/28

GOES Flare Class: *M1.3*
Flare Location (degrees):
-19, -66

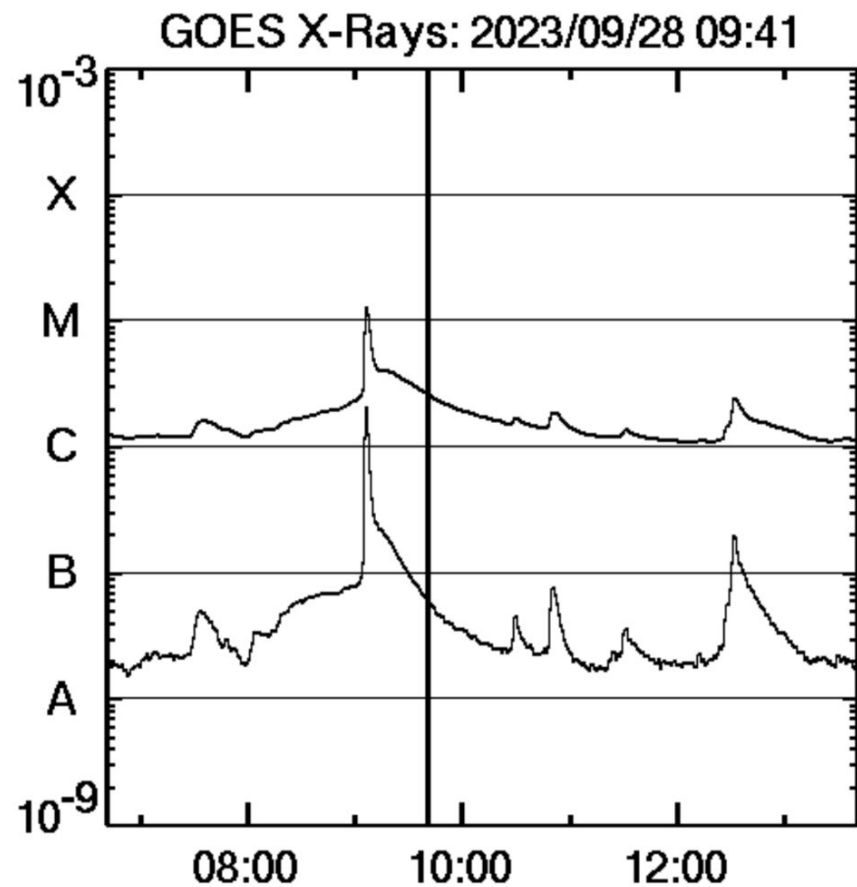
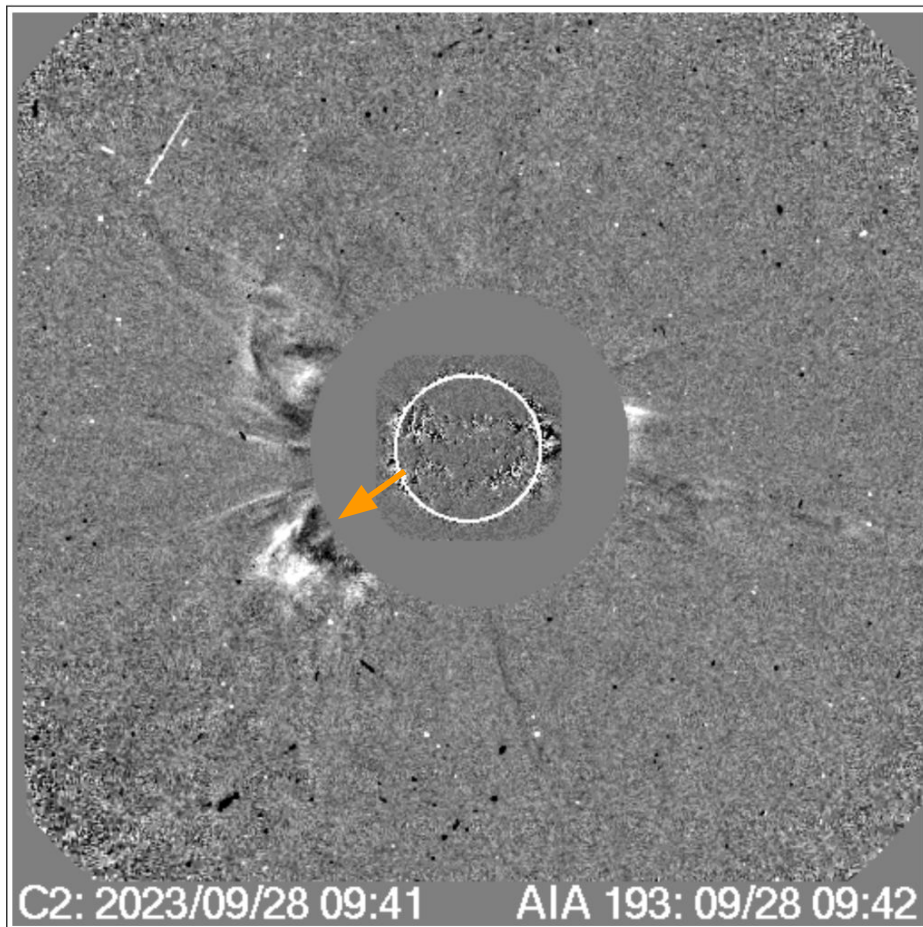


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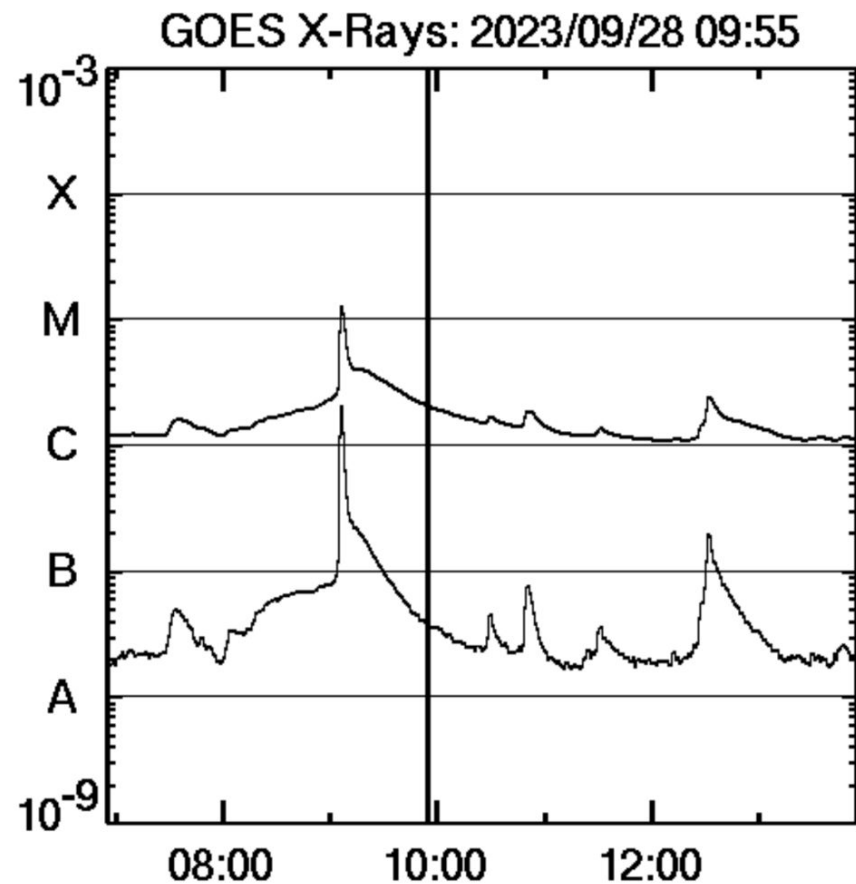
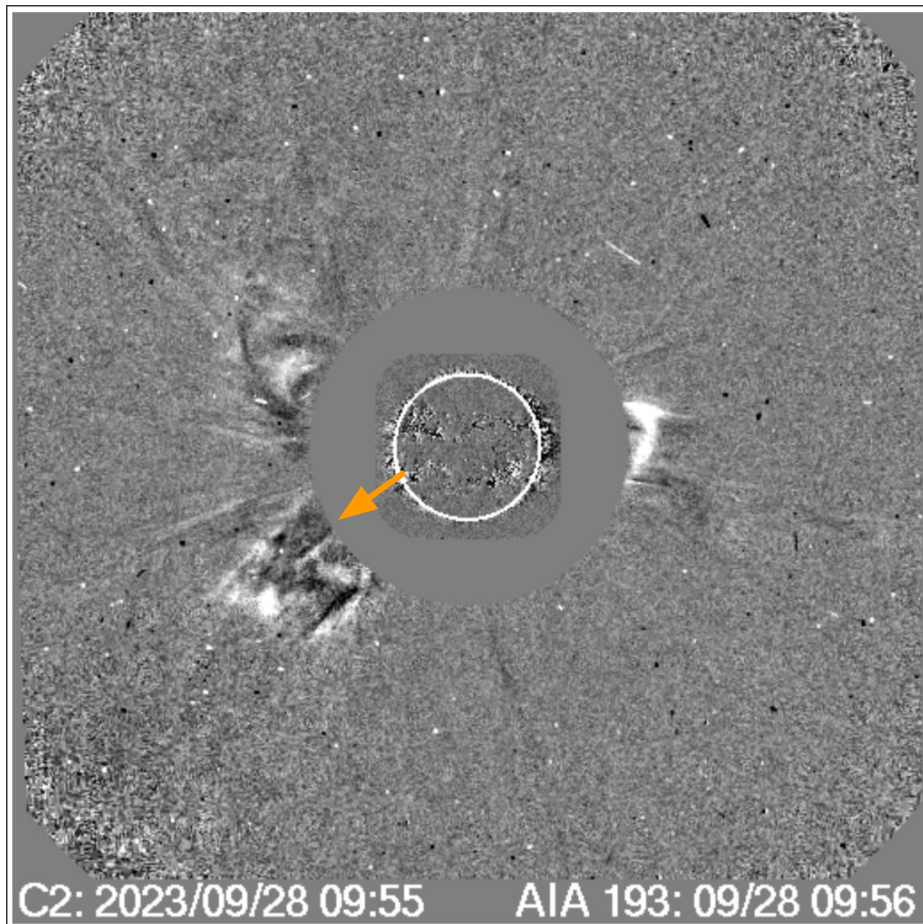
LASCO observations 2023/09/28

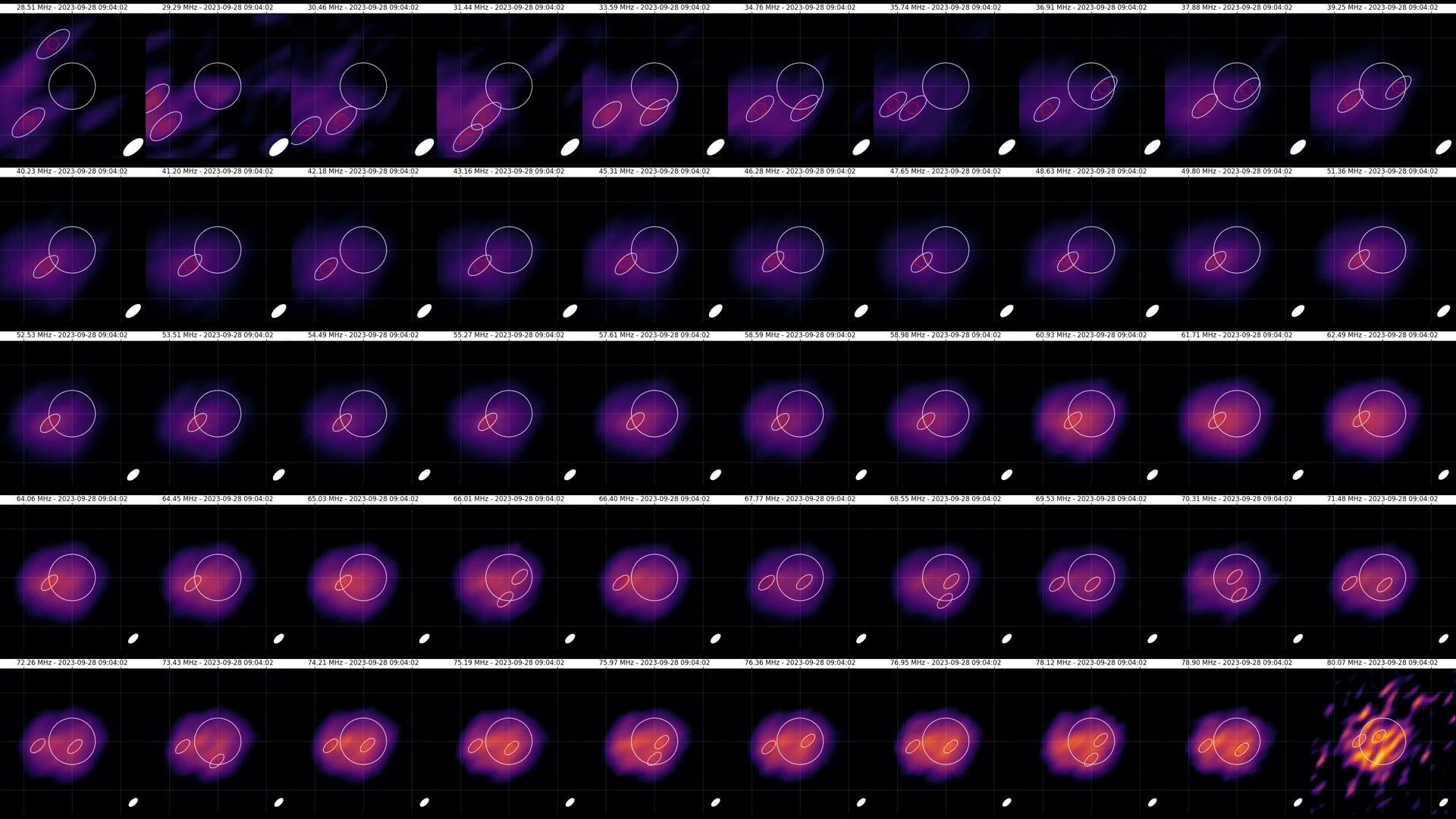


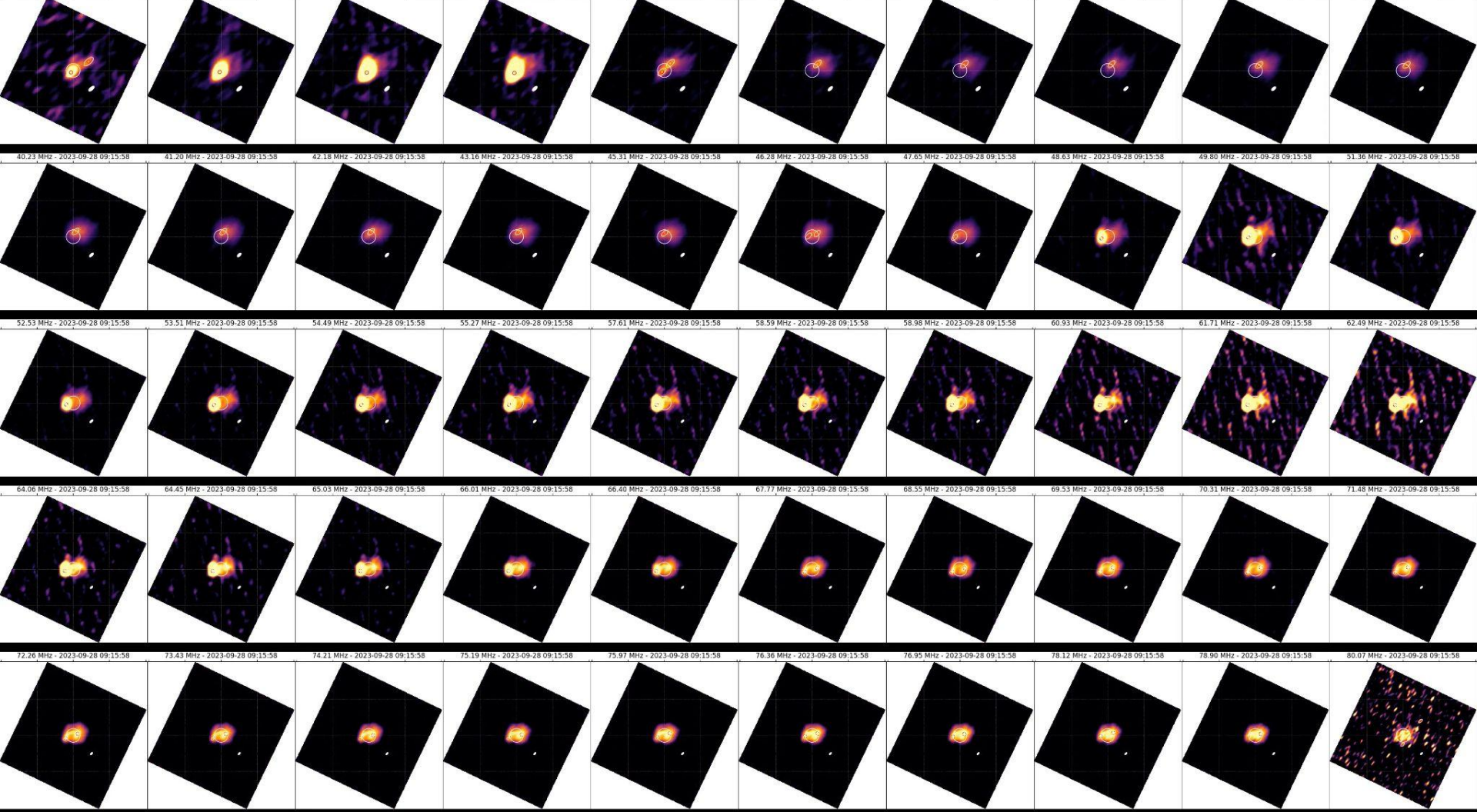
LASCO observations 2023/09/28



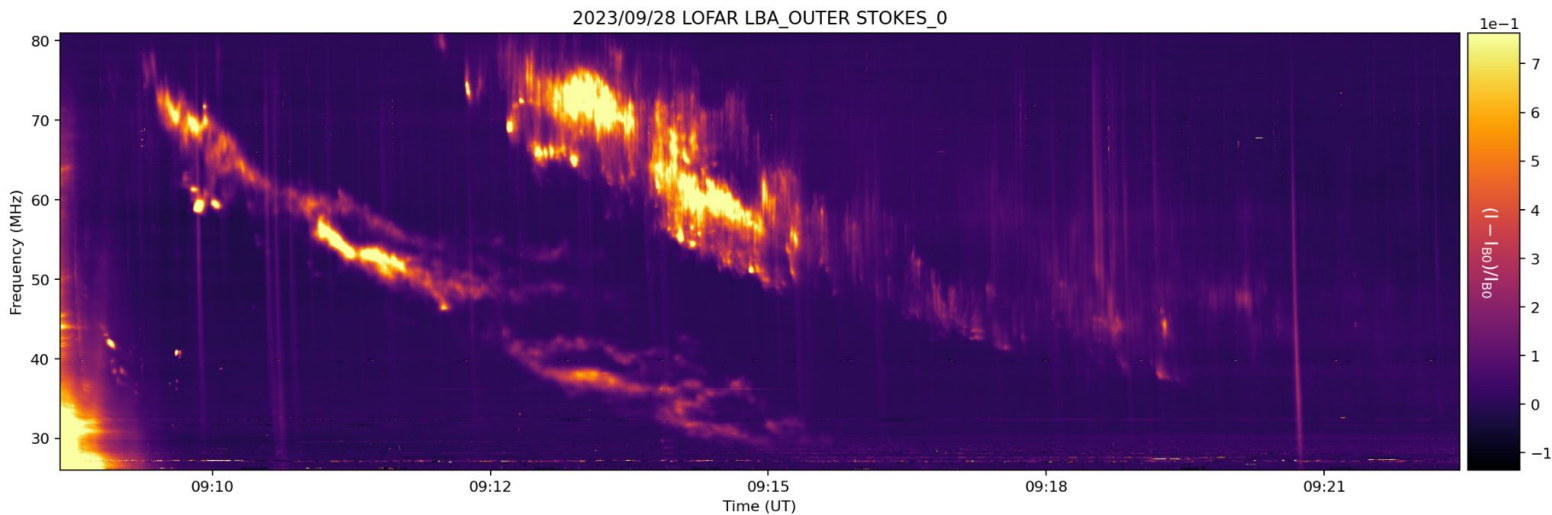
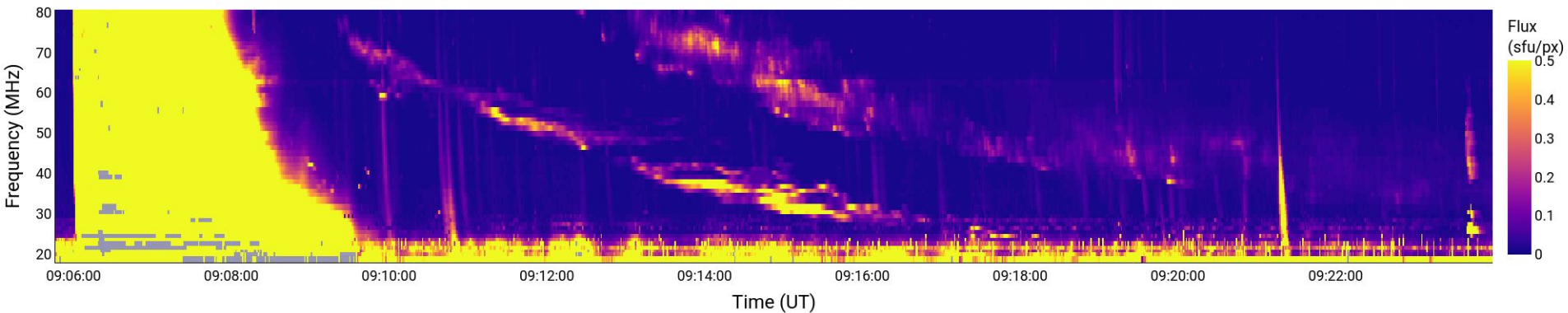
LASCO observations 2023/09/28



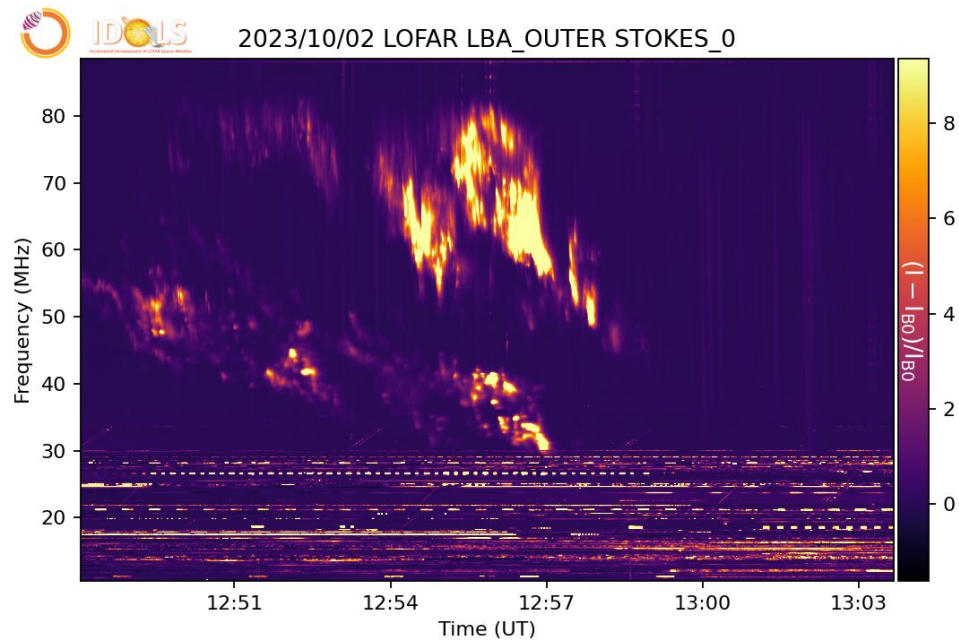
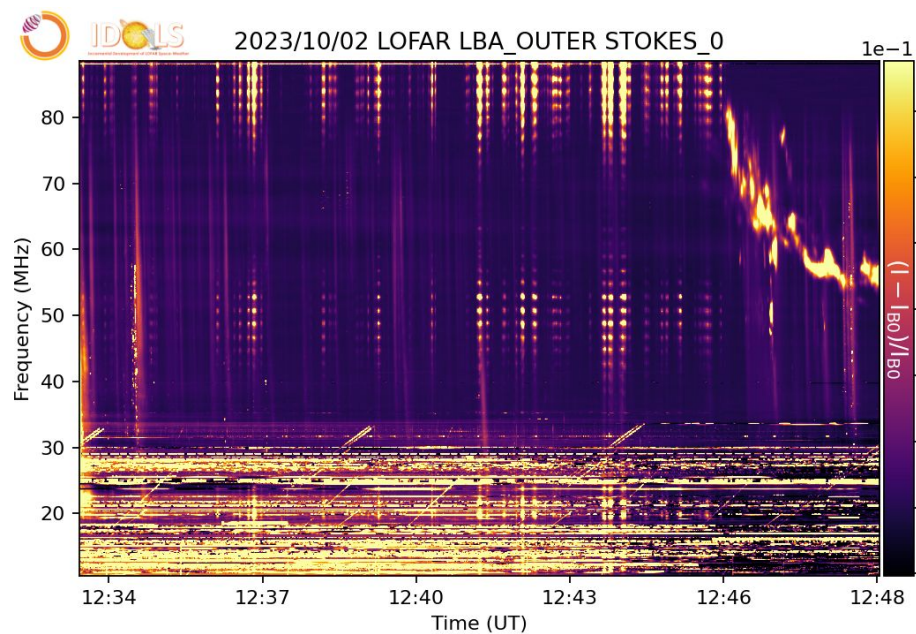


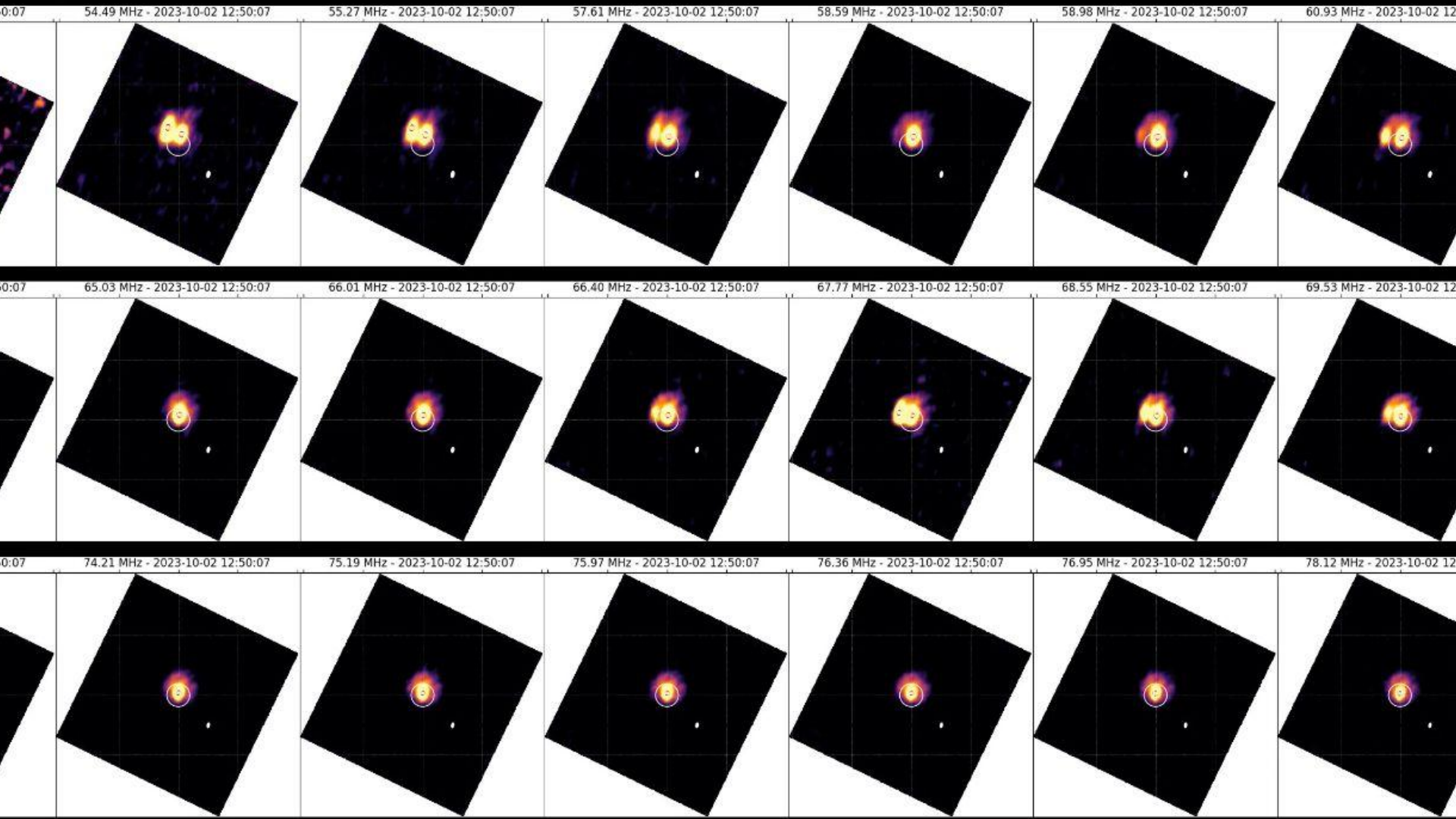


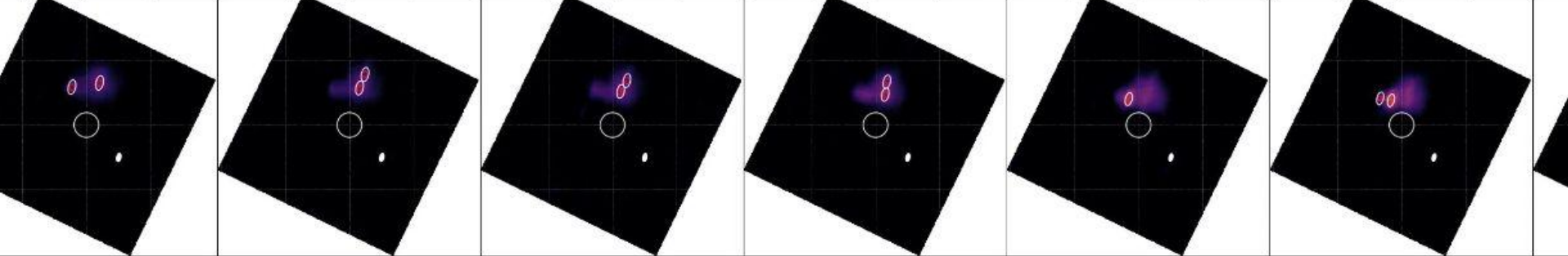
2024/10/02 12:45-13:05 Type II burst source extraction



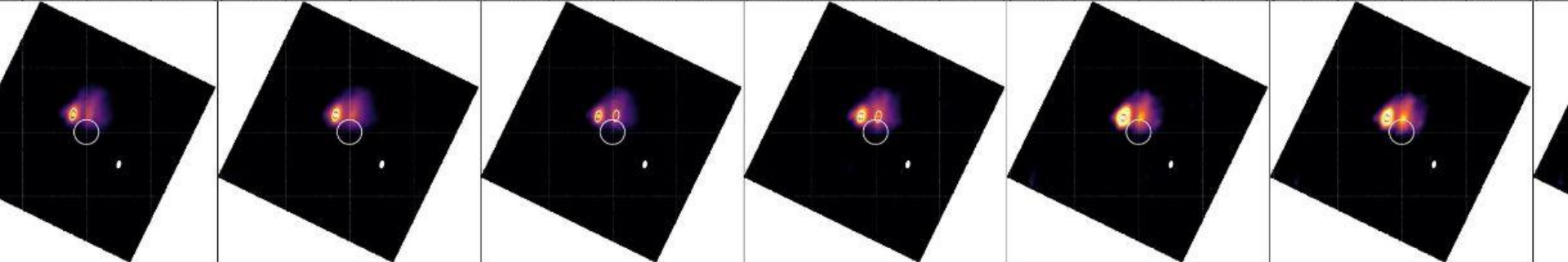
2024/10/02 12:45-13:05 Type II burst



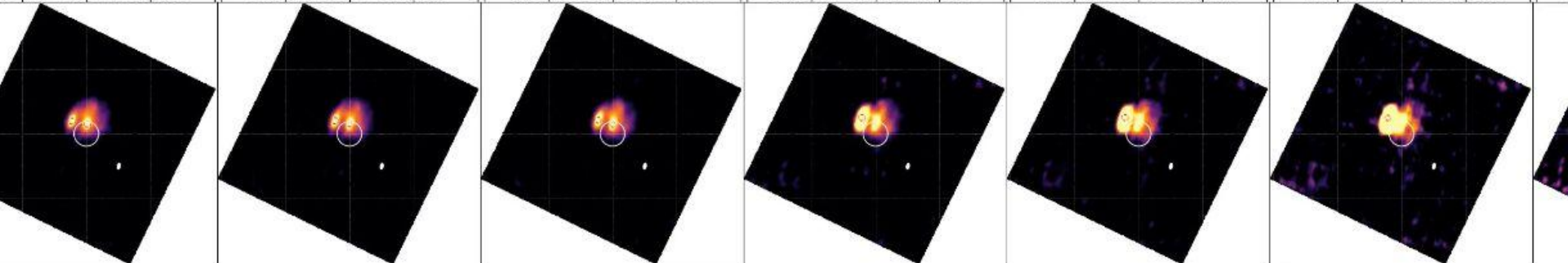




53 MHz - 2023-10-02 12:46:26 53.51 MHz - 2023-10-02 12:46:26 54.49 MHz - 2023-10-02 12:46:26 55.27 MHz - 2023-10-02 12:46:26 57.61 MHz - 2023-10-02 12:46:26 58.59 MHz - 2023-10-02 12:46:26



60.06 MHz - 2023-10-02 12:46:26 64.45 MHz - 2023-10-02 12:46:26 65.03 MHz - 2023-10-02 12:46:26 66.01 MHz - 2023-10-02 12:46:26 66.40 MHz - 2023-10-02 12:46:26 67.77 MHz - 2023-10-02 12:46:26

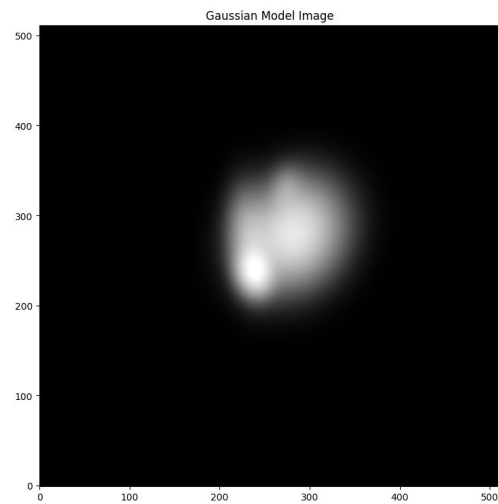
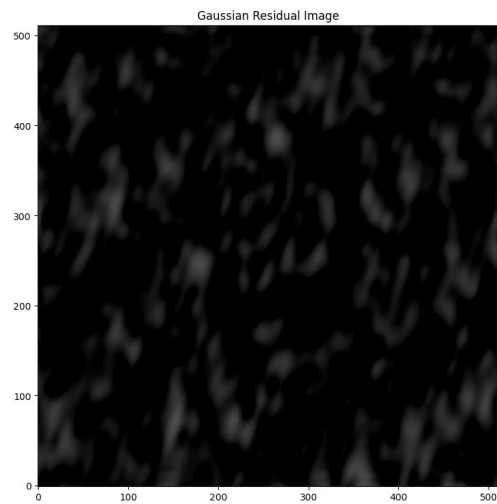
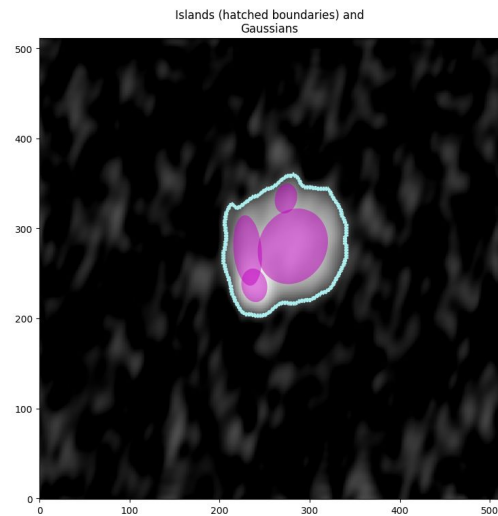
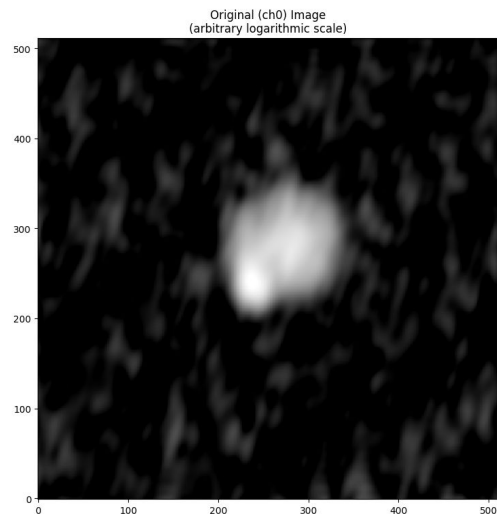


73.43 MHz - 2023-10-02 12:46:26 74.21 MHz - 2023-10-02 12:46:26 75.19 MHz - 2023-10-02 12:46:26 75.97 MHz - 2023-10-02 12:46:26 76.36 MHz - 2023-10-02 12:46:26

PyBDSF source extraction from interferometry snapshots

Challenges:

- Source tracking
- Stability
- Processing time



2023/09 24th vs 28th Comparison

- Similar solar context
- Significant ionospheric refraction on the 24th
- Different sources dynamics observed in interferometry
- LoG source extraction implementation useful only for brighter emissions
- Main dynamic region associated with type II in both
- Additional activity present during the events

Team effort

- Kamen Kozarev
- Pietro Zucca
- Mattia Mancini
- Mohamed Nedal



DISCOVERER